

Comparison of different km³ designs using Antares tools

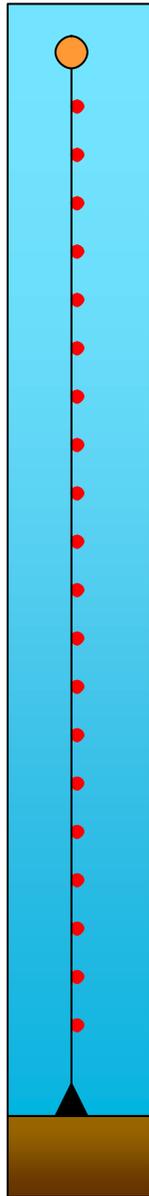
D. Zaborov (ITEP, Moscow)

- Three kinds of detector geometry
- Incoming muons within 1 .. 1000 TeV energy range
- Detector efficiency and angular resolution obtained with the Antares tools

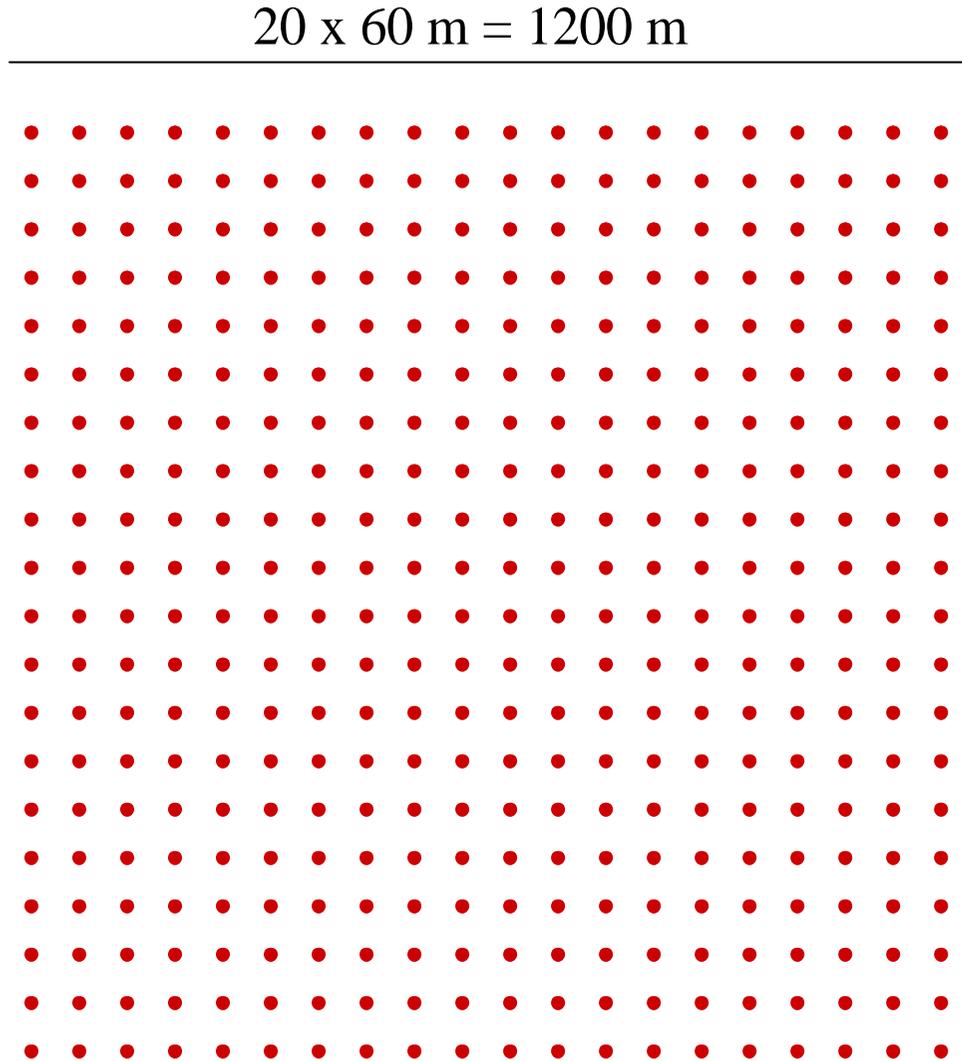
A large homogeneous KM3 detector (8000 PMTs)

Structure of
the string

20 x 60 m = 1200 m

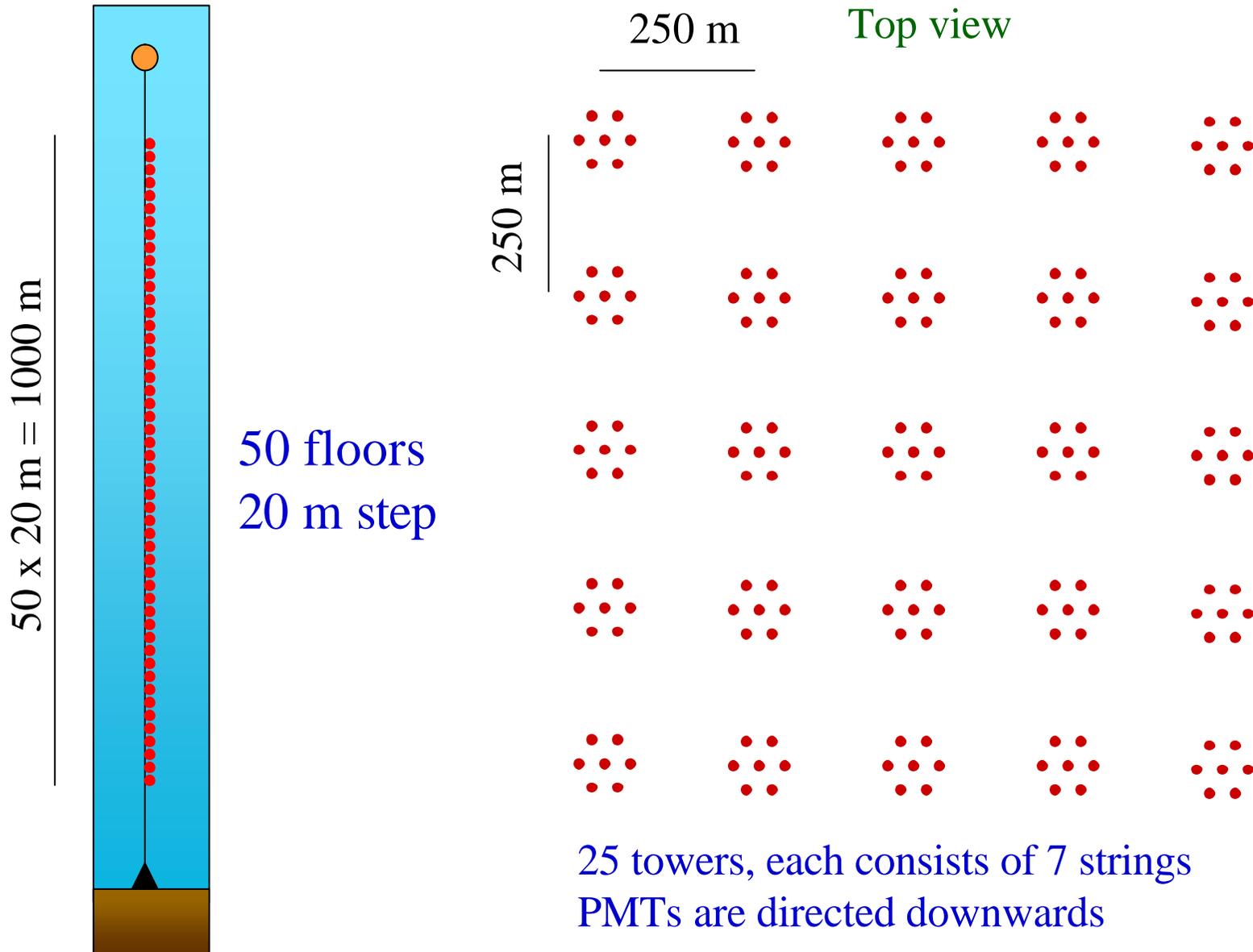


20 x 60 m = 1200 m

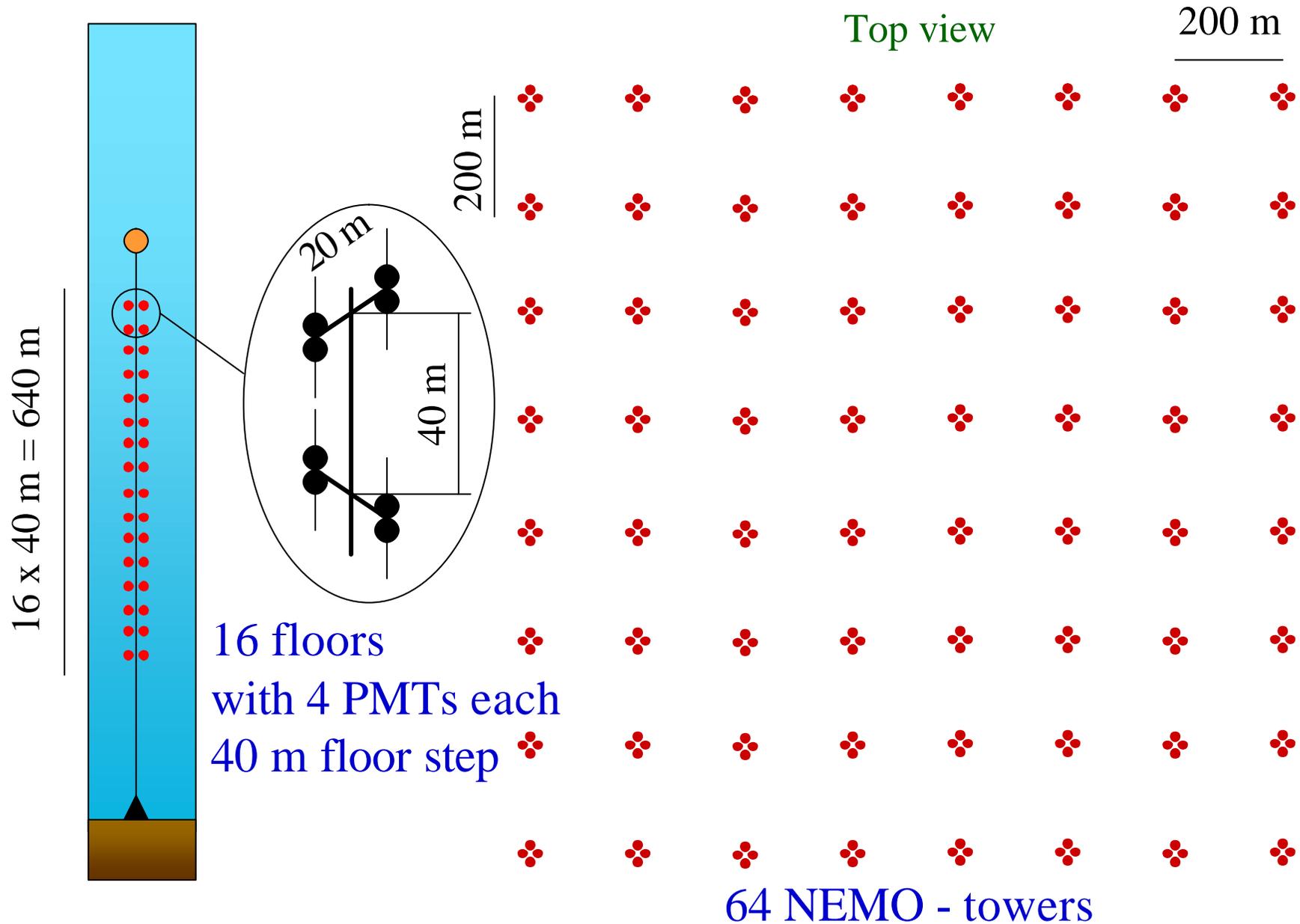


homogeneous lattice 20 x 20 x 20 downward-looking
10 “ photomultiplier tubes

A large NESTOR – like detector (8750 PMTs)



A large NEMO – like detector (4096 PMTs)



Detector geometries: Common features

Set 1:

PMTs: 10 inch

Characteristics from Antares Hamamatsu R7820 used
(eff. Area = 0.44 cm^2 , angular response, QE)

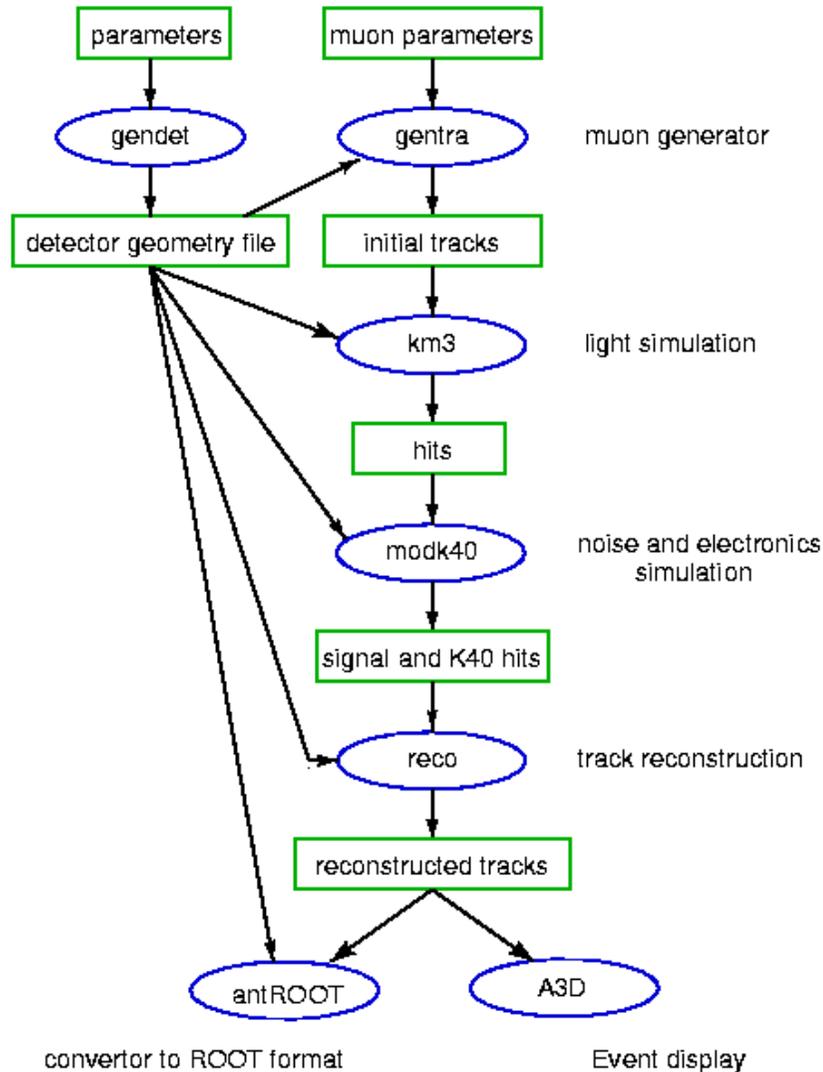
Set 2:

15 inch PMTs: same characteristics but eff. Area = $(15/10)^2$

No PMT multiplotts used

i.e. Tight coincidences between close PMTs not possible
replaced by high pulse condition (above 2.5 , 3.5 pe)

Event Sample



Energy range: 1TeV - 1PeV

Spectrum: $1/E$ (flat in $\log(E)$)

lower hemisphere **isotropic**

Surface drawing

km3 with **partic-0.0075**

60 (135) kHz noise light

Simplified **digitisation**

(2 ARS 25nsec integration)

recov4r2 with **AartStrategy**

Analysis **antroot** and **a3d**

Program modifications

Most of the code was easily scalable for use with 1km³ detector

Some parameters in include files had to be changed like:

(km3.inc, phomul.inc)

max. number of PMTs, clusters, strings

Patch work in km3 to allow use of 15 inch PMT without creating new diffusion tables

Reco: modification of Select and Filter routines to

adapt to high pulse condition and absence of coincidences

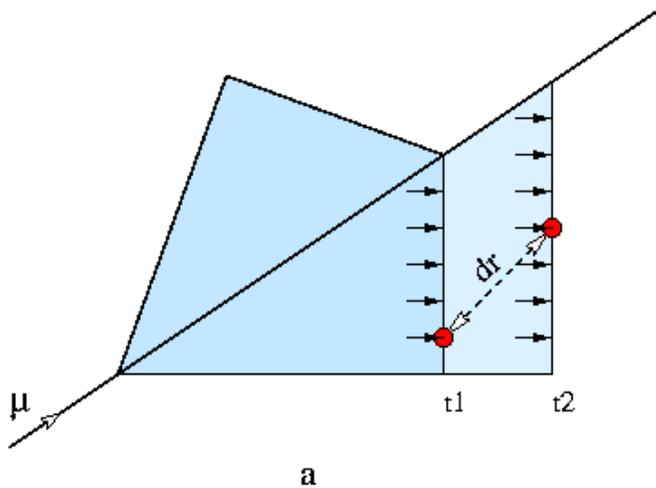
(2.5 pe for 10 inch PMT, 3.5pe for 15inch PMT)

Modified causality filter

Usual filter:

$$\text{abs}(dt) < dr/v_{\text{light}}$$

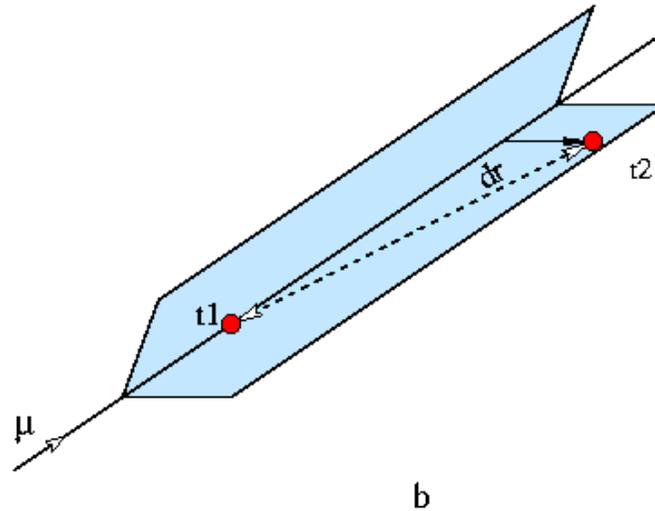
Plane wave approximation
 $dr \ll x$ (photon travel path)



Large detectors (new condition):

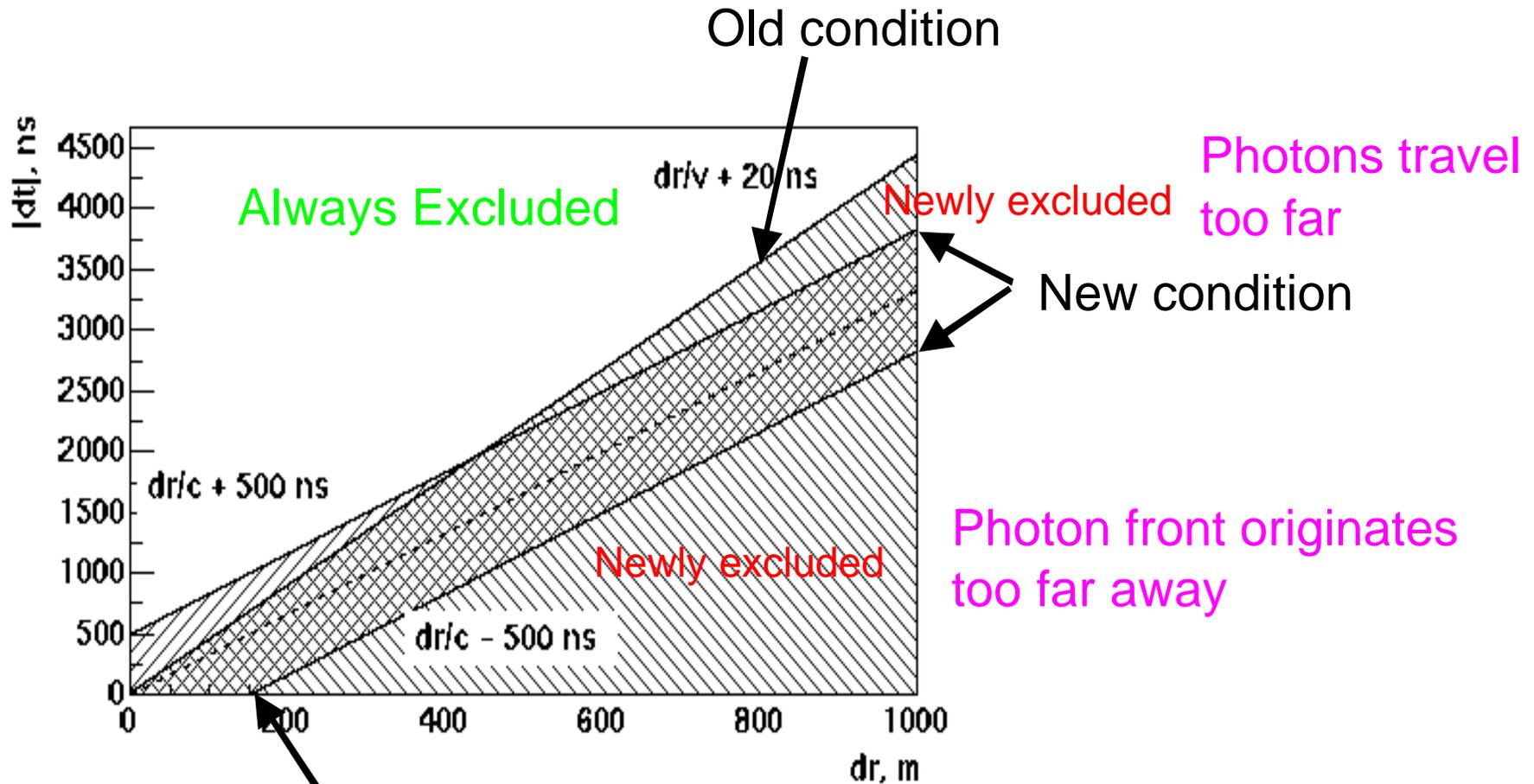
$$\text{abs}[\text{abs}(dt) - dr/c_{\text{light}}] < 500\text{nsec}$$

Approximation $x \ll dr$
(hits close to muon track,
takes into account absorption)



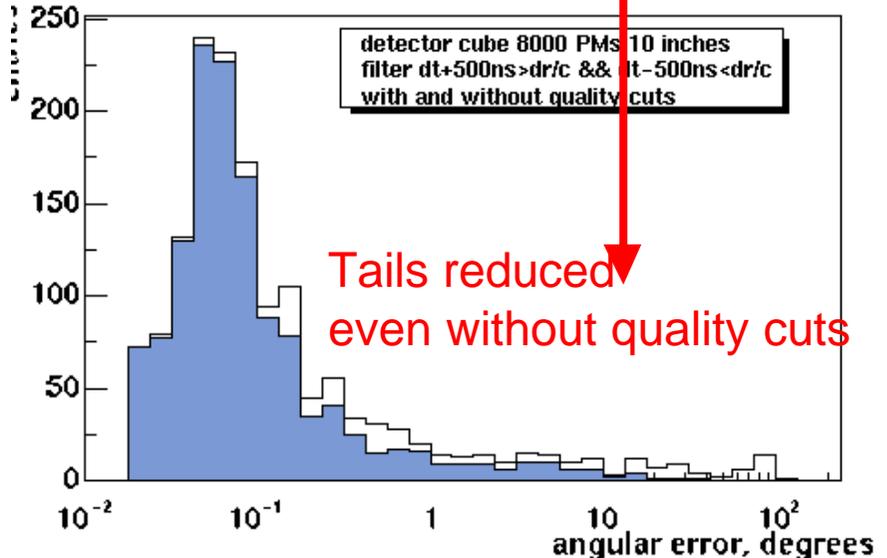
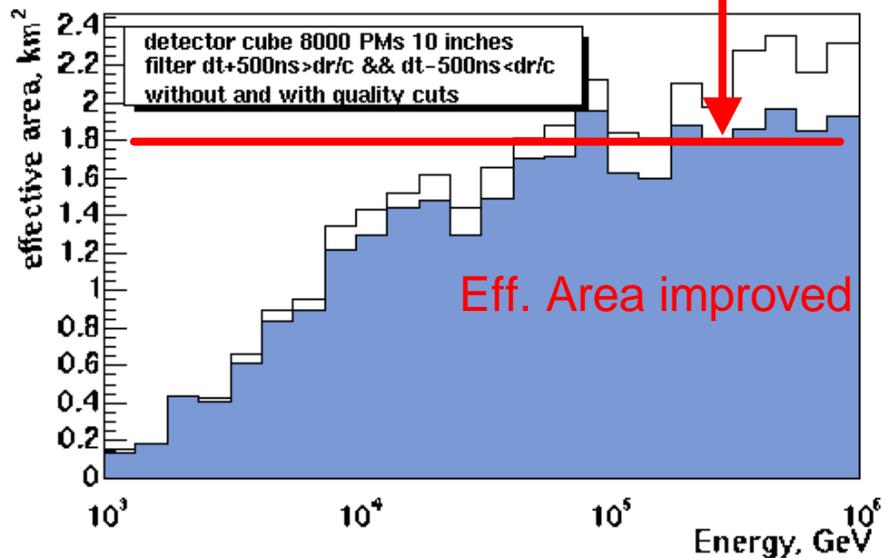
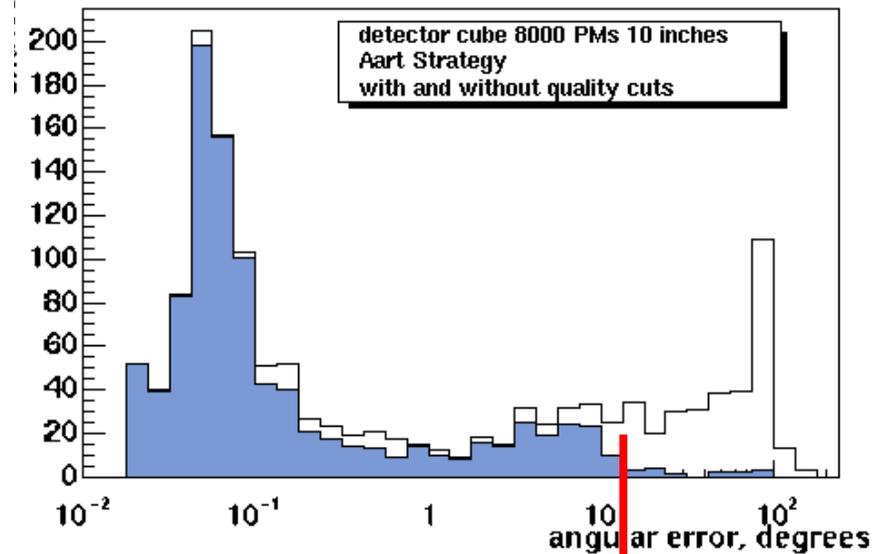
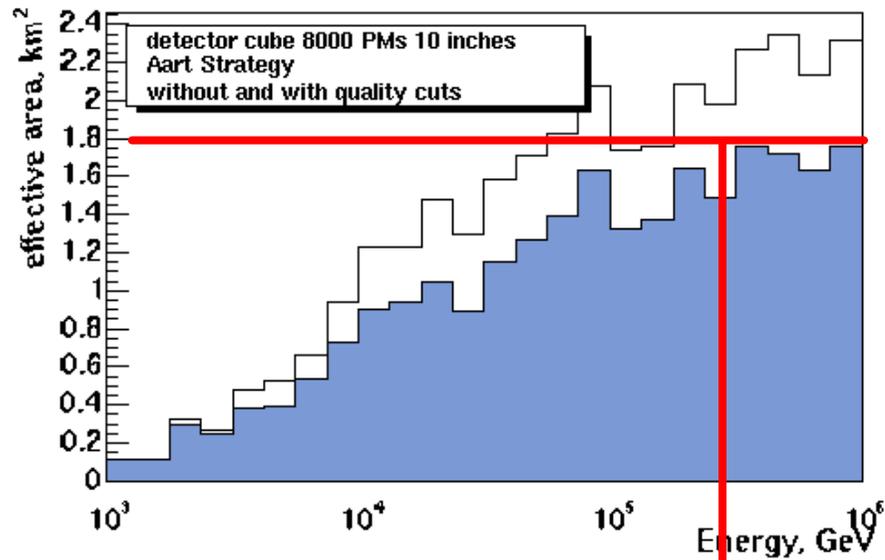
Modified causality filter

Logical AND between both conditions



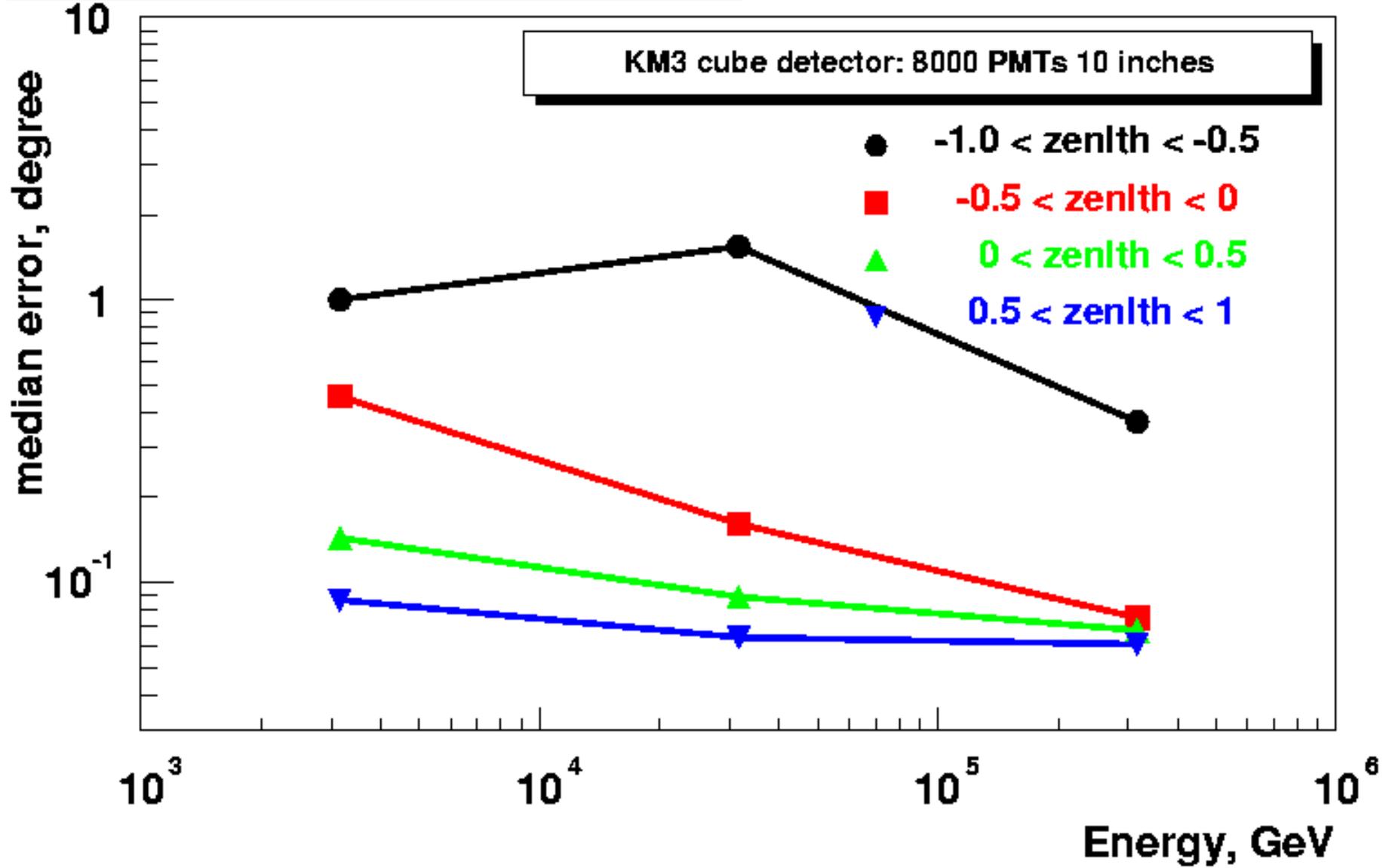
Effective for $dr > 150\text{m}$, probably negligible for 0.1km^2 detector

Effect of new causality filter

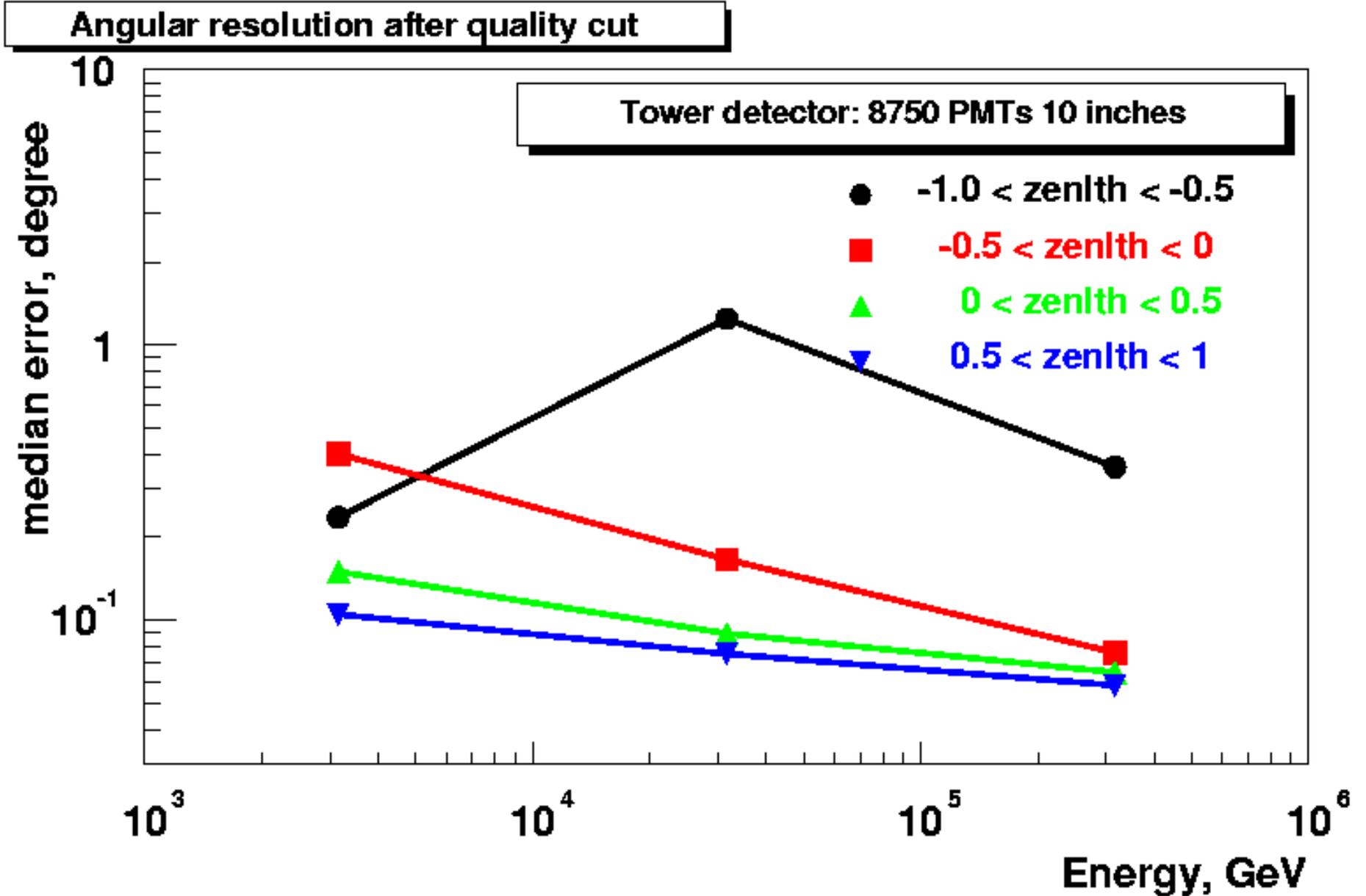


Angular resolution of the homogeneous detector

Angular resolution after quality cut

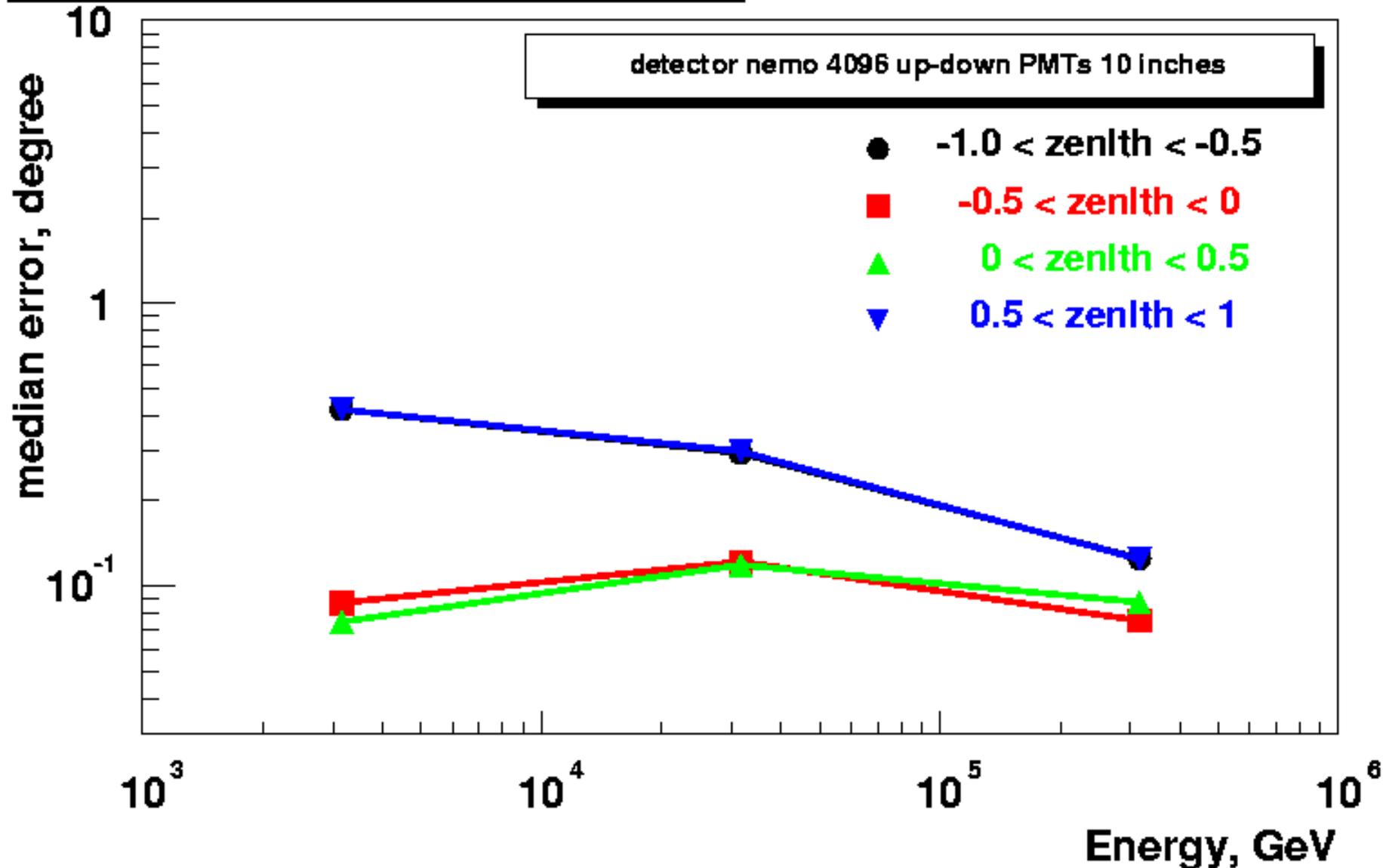


Angular resolution of the NESTOR-like detector



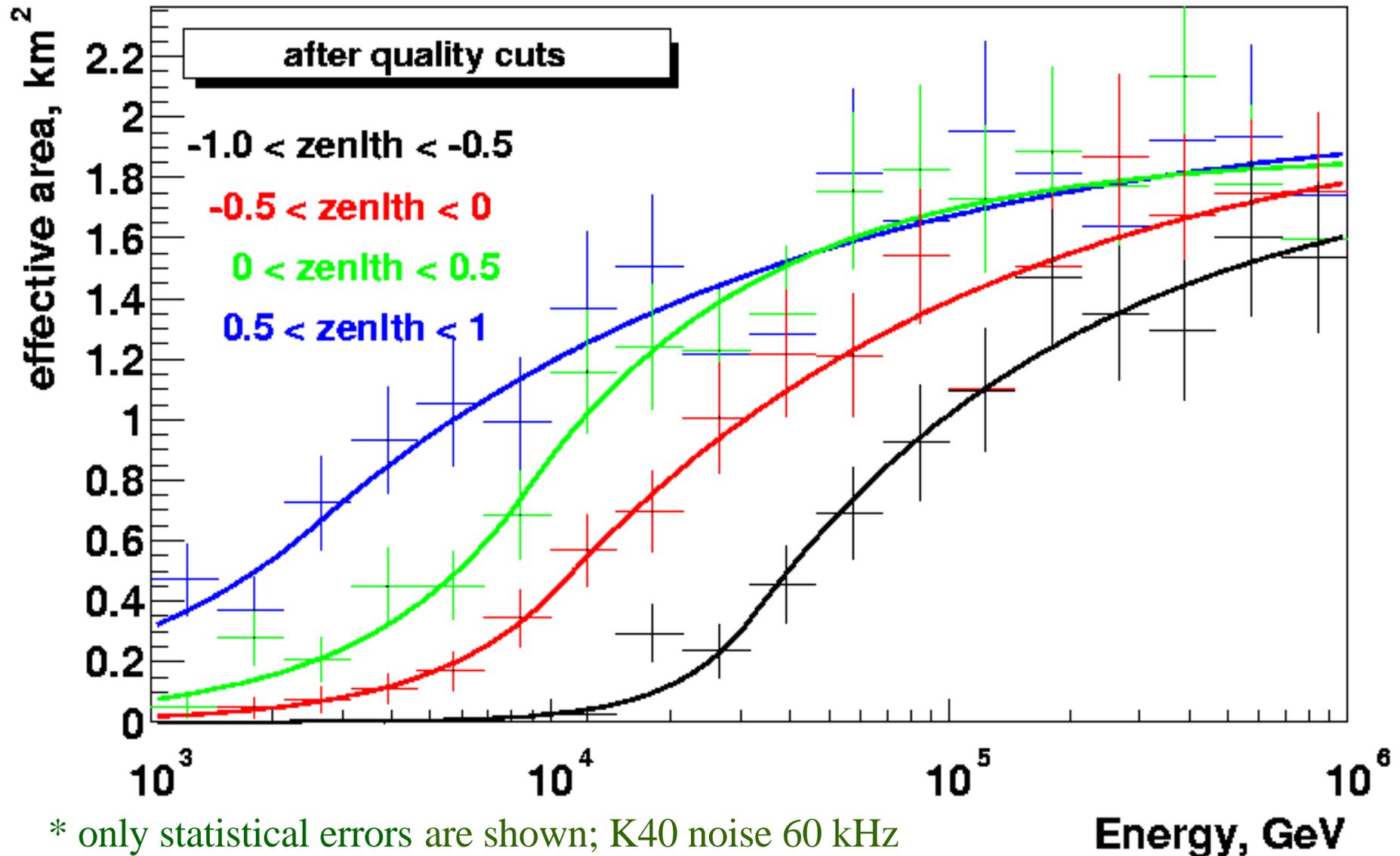
Angular resolution of the NEMO-like detector

Angular resolution after quality cut



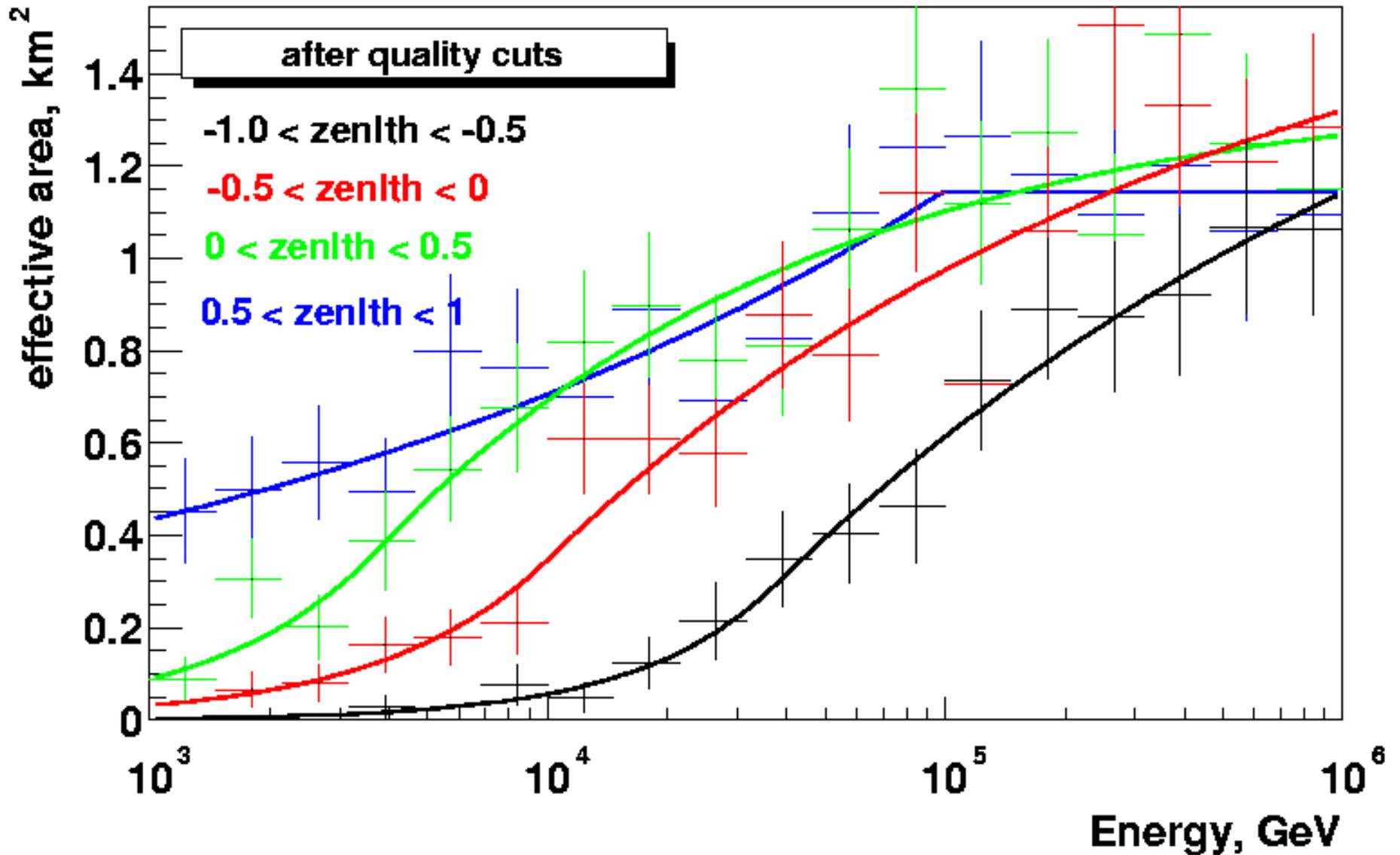
Effective area of the homogeneous detector for muons

KM3 cube detector: 8000 PMTs 10 inches



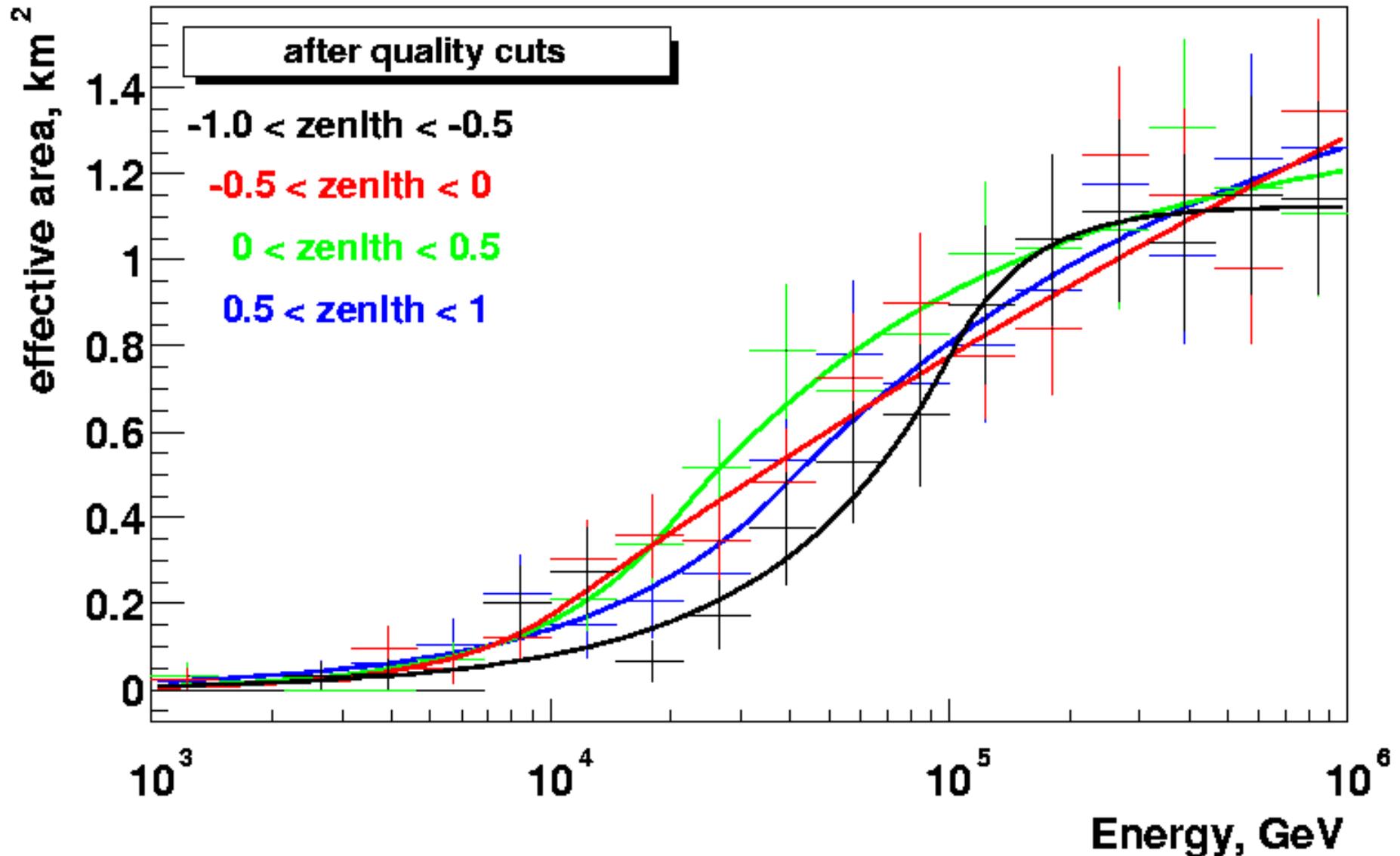
Effective area of the NESTOR-like detector for muons

Tower detector: 8750 PMTs 10 inches



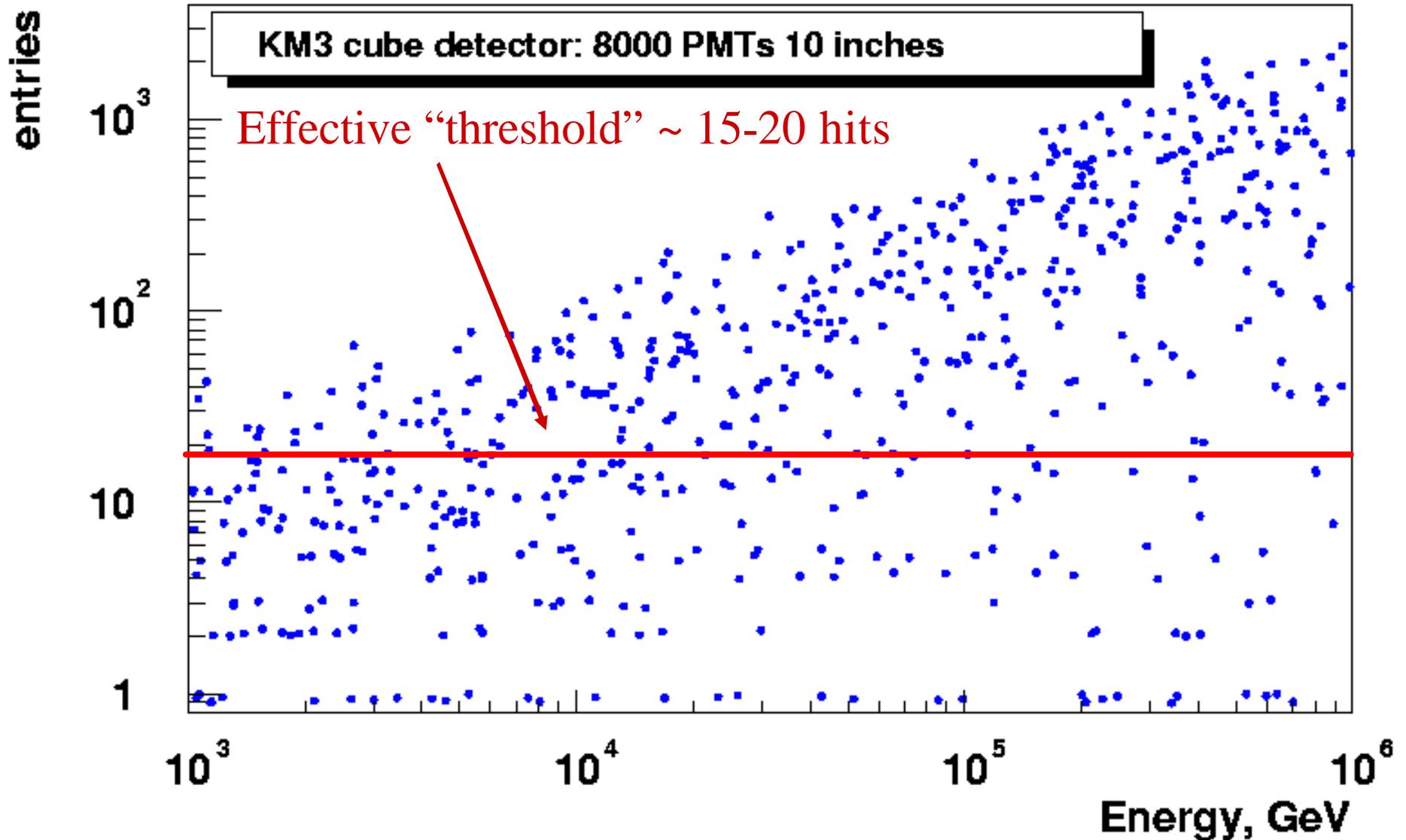
Effective area of the NEMO-like detector for muons

detector nemo 4096 up-down PMTs 10 inches



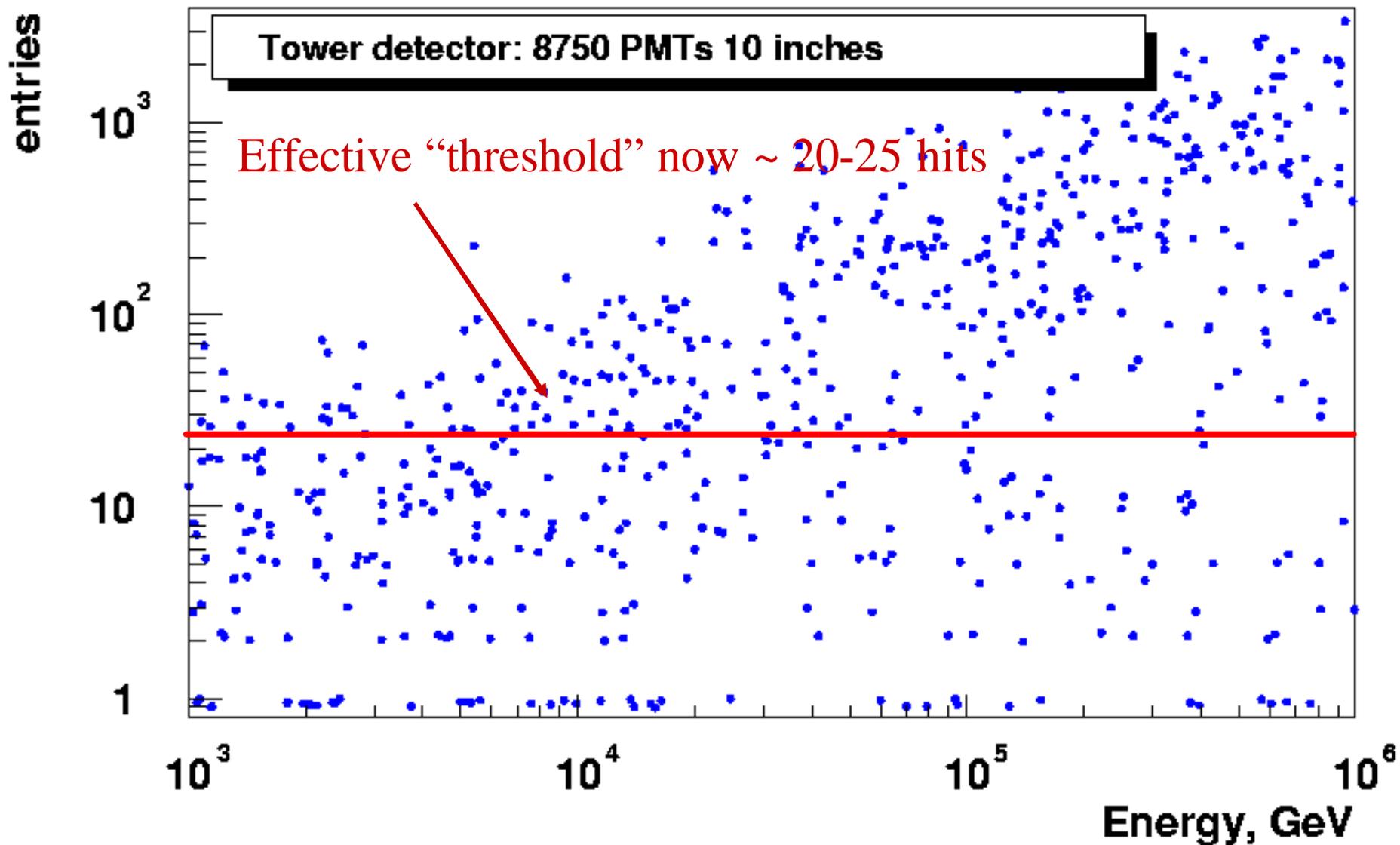
Homogeneous geometry: number of detected hits vs. muon energy

number of signal hits per event



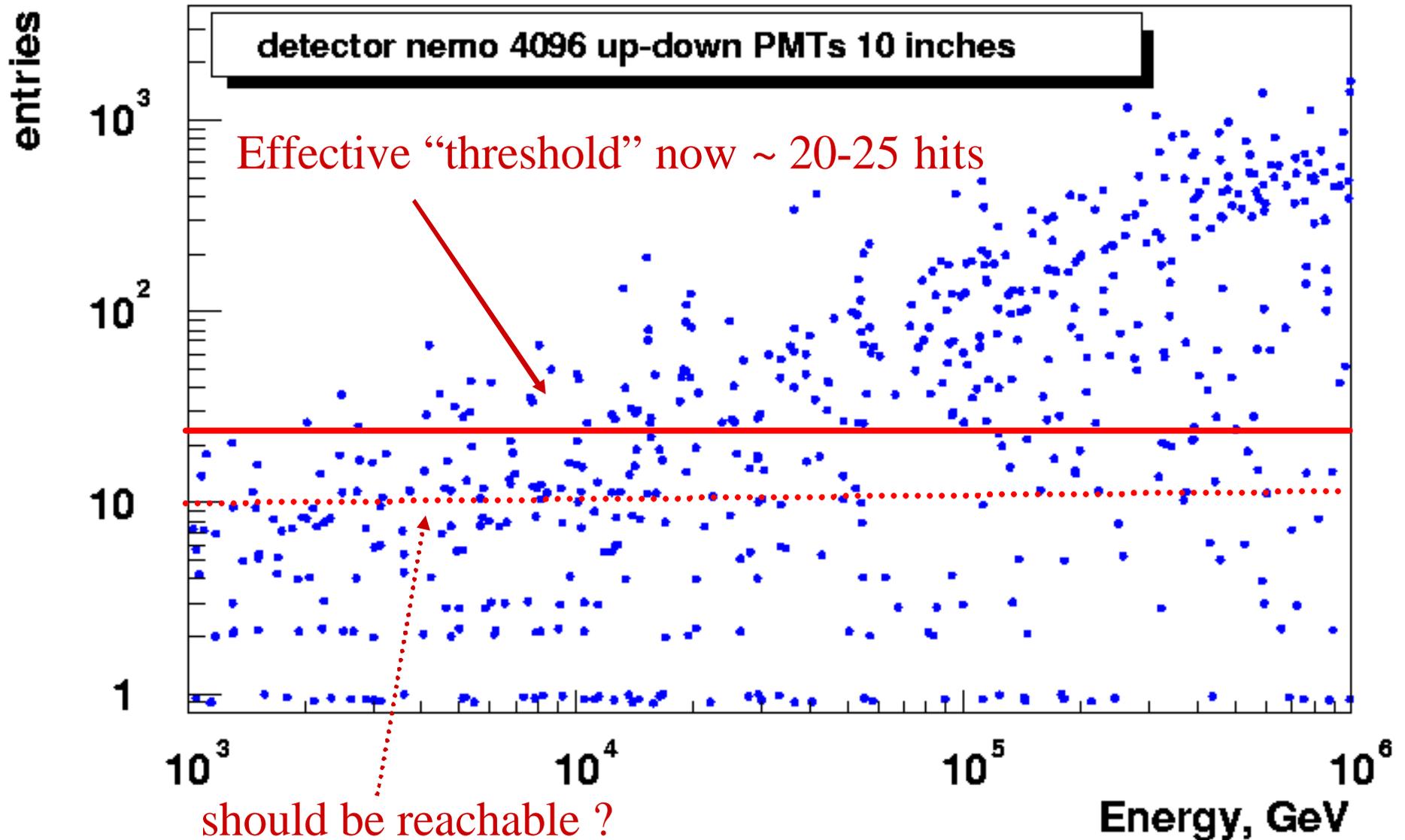
NESTOR-like geometry: number of detected hits vs. muon energy

number of signal hits per event

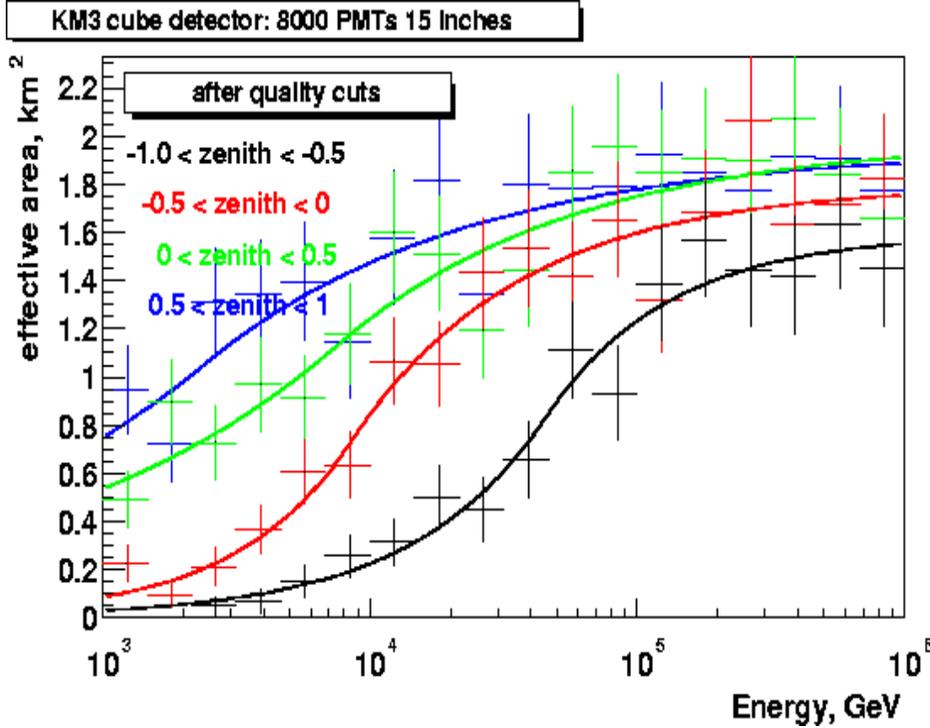
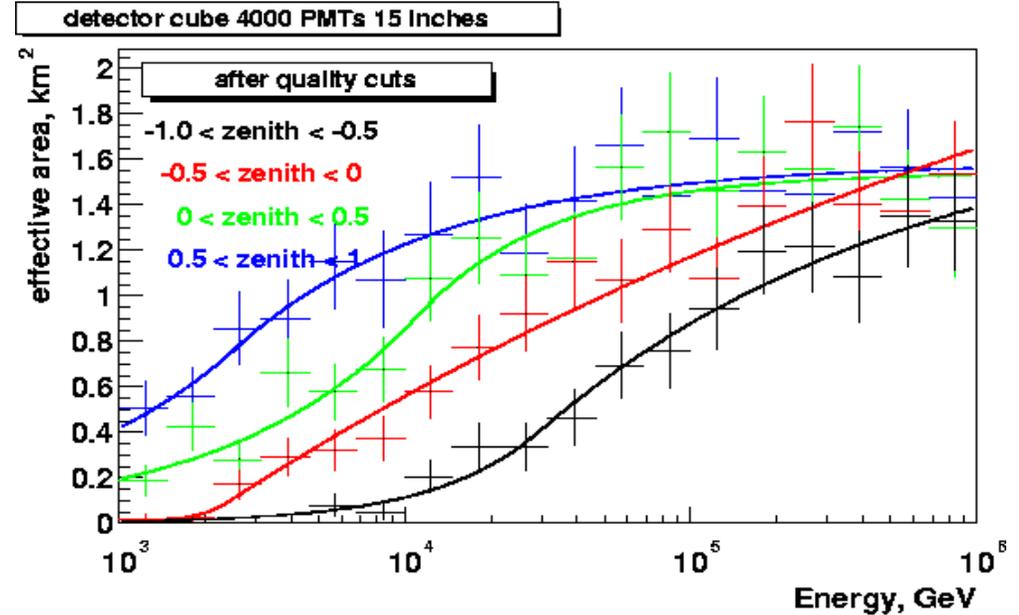
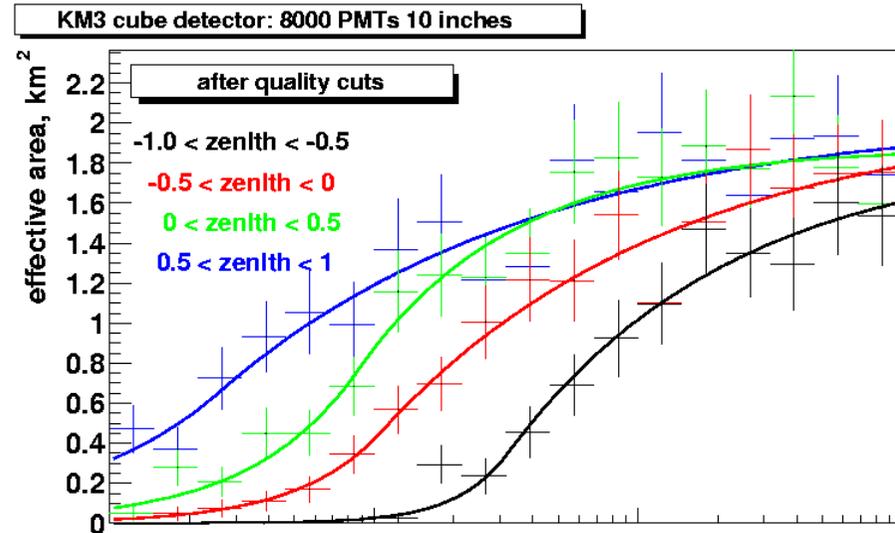


NEMO geometry: number of detected hits vs. muon energy

number of signal hits per event



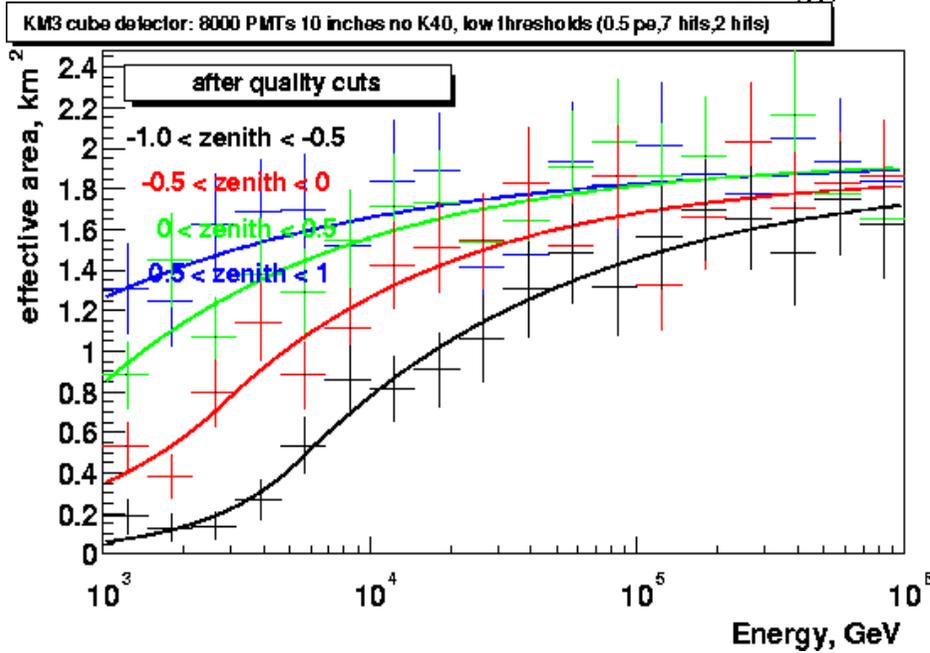
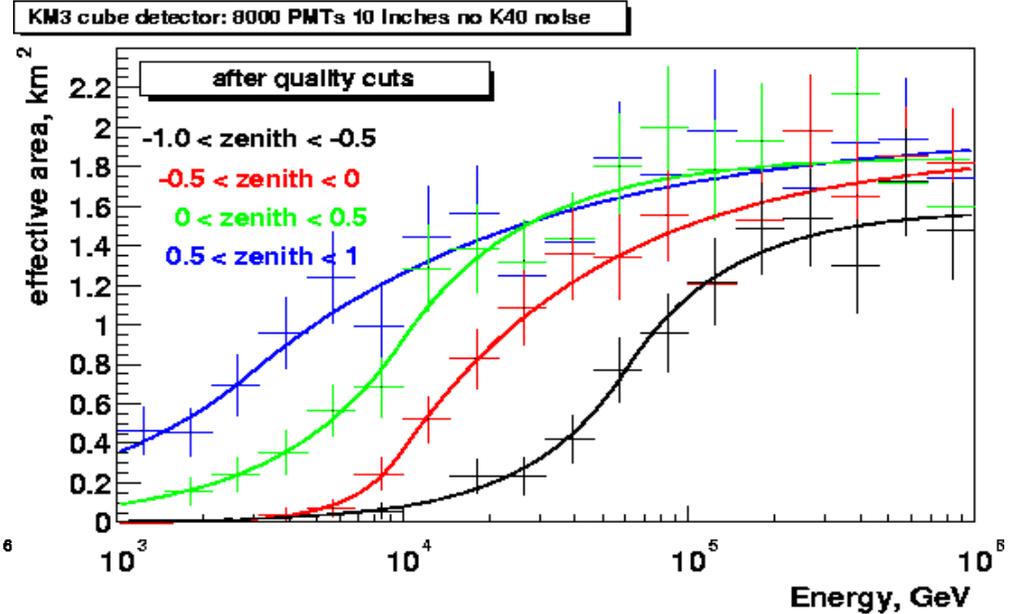
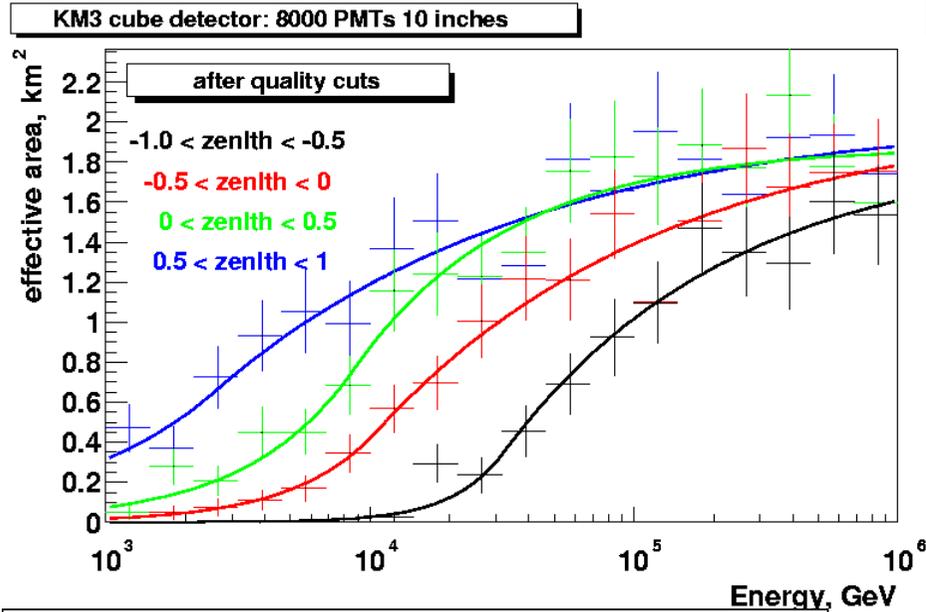
Effective area dependence on PMT size



At low energies dependence
on total photocathode area

At high energies
dependence on PMT number

Effective area dependence on noise



Used reconstruction filters
noise hits effectively

At low energies
lowering thresholds with
low noise improves
performance

Conclusions

- Monte-Carlo simulation of Cherenkov light production and muon track reconstruction have been performed using Antares tools.
- The homogeneous geometry demonstrates high detector efficiency starting from 1 TeV.
- Effective area reached with the NESTOR-like geometry is smaller than obtained with the homogeneous geometry.
- The NEMO-like detector has a significantly smaller effective area than the homogeneous configuration mainly due to its lower PMT density.
- All the geometries show angular resolution in muon track reconstruction of about 0.1 degree or better.
- High requirements of the reconstruction software for the number of hits, especially with non-homogeneous detectors, have been discovered and require further investigations.