

NESTOR

Neutrino **E**xtended **S**ubmarine **T**elescope with **O**ceanographic **R**esearch

The NESTOR O.M.

The optical module for the NESTOR neutrino telescope

E.G. Anassontzis^a, P. Ioannou^a, S. Katsanevas^a, Ch. Kourkouvelis^a, J. McNutt^a,
A. Manousakis-Katsikakis^a, L. Moraitis^a, L.K. Resvanis^a, S.A. Sotiriou^{a,*},
V. Tsagli^a, G. Voulgaris^a, I. Siotis^a, G. Fanourakis^a, G. Grammatikakis^a,
A.E. Ball^a, S. Bottai^a, A. Cartacci^a, B. Monteleoni^a, U. Keusen^a, P. Koske^a,
V.A. Zhukov^a, V.K. Rucol^a, V.V. Ledenev^a

^a Physics Laboratory, University of Athens, Greece

^b NCSR "DEMOKRITOS", Greece

^c University of Crete, Greece

^d CERN, Switzerland

^e University of Florence, Italy

^f University of Kiel, Germany

^g Institute for Nuclear Research, Russian Academy of Sciences, Russia

^h Institute of Oceanology, Russian Academy of Sciences, Russia

ⁱ Experimental Design Bureau of Oceanological Equipment, Russian Academy of Sciences, Russia

The NESTOR collaboration

Received 24 October 2000; received in revised form 12 February 2001; accepted 19 March 2001

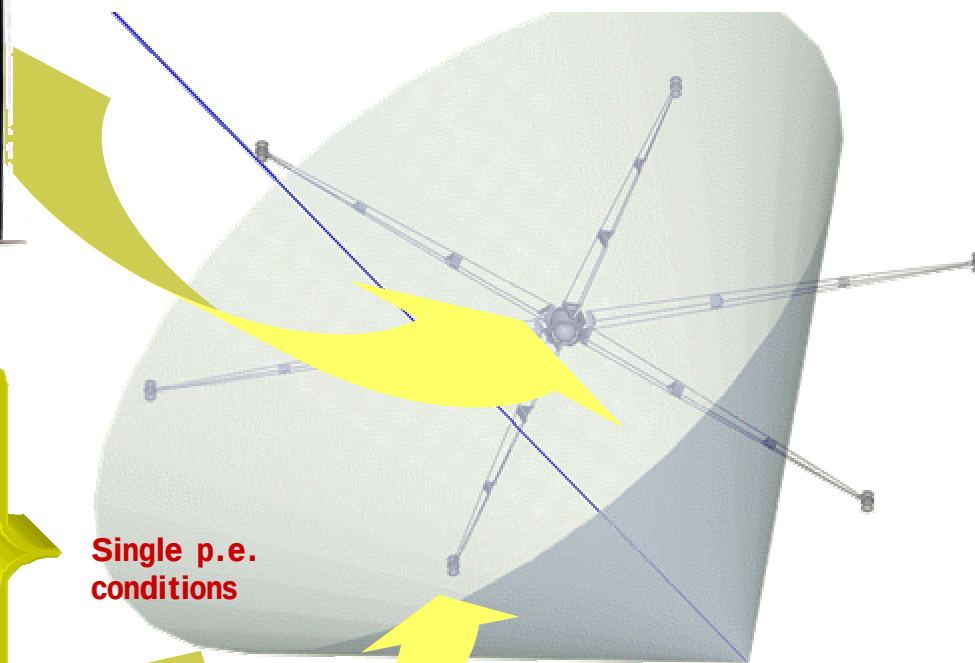
Abstract

NESTOR is a deep-sea Cherenkov neutrino detector now under construction for deployment in the Mediterranean off Greece. Its key component is an optical module employing a photomultiplier tube with a 15 in. hemispherical photocathode in a transparent glass pressure housing. Extensive tests have been made on the sensitivity, uniformity, time resolution, noise rates and mechanical properties of the module: several test deployments have been made at sea. © 2002 Elsevier Science B.V. All rights reserved.

Keywords: Neutrino telescope; Cherenkov detector; Optical sensor

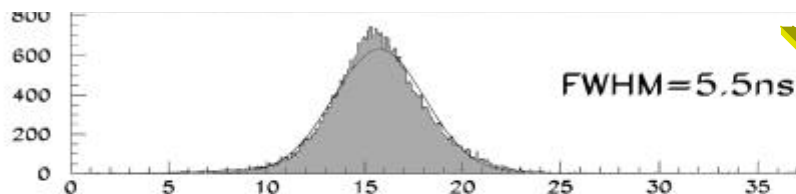


The Detector

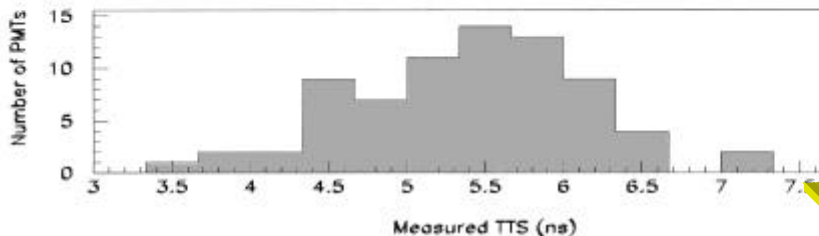


Single p.e.
conditions

Potentiometer

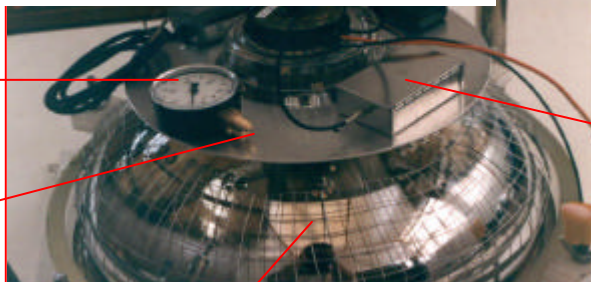


TTS at single p.e. conditions



Pressure
gauges

Al disk

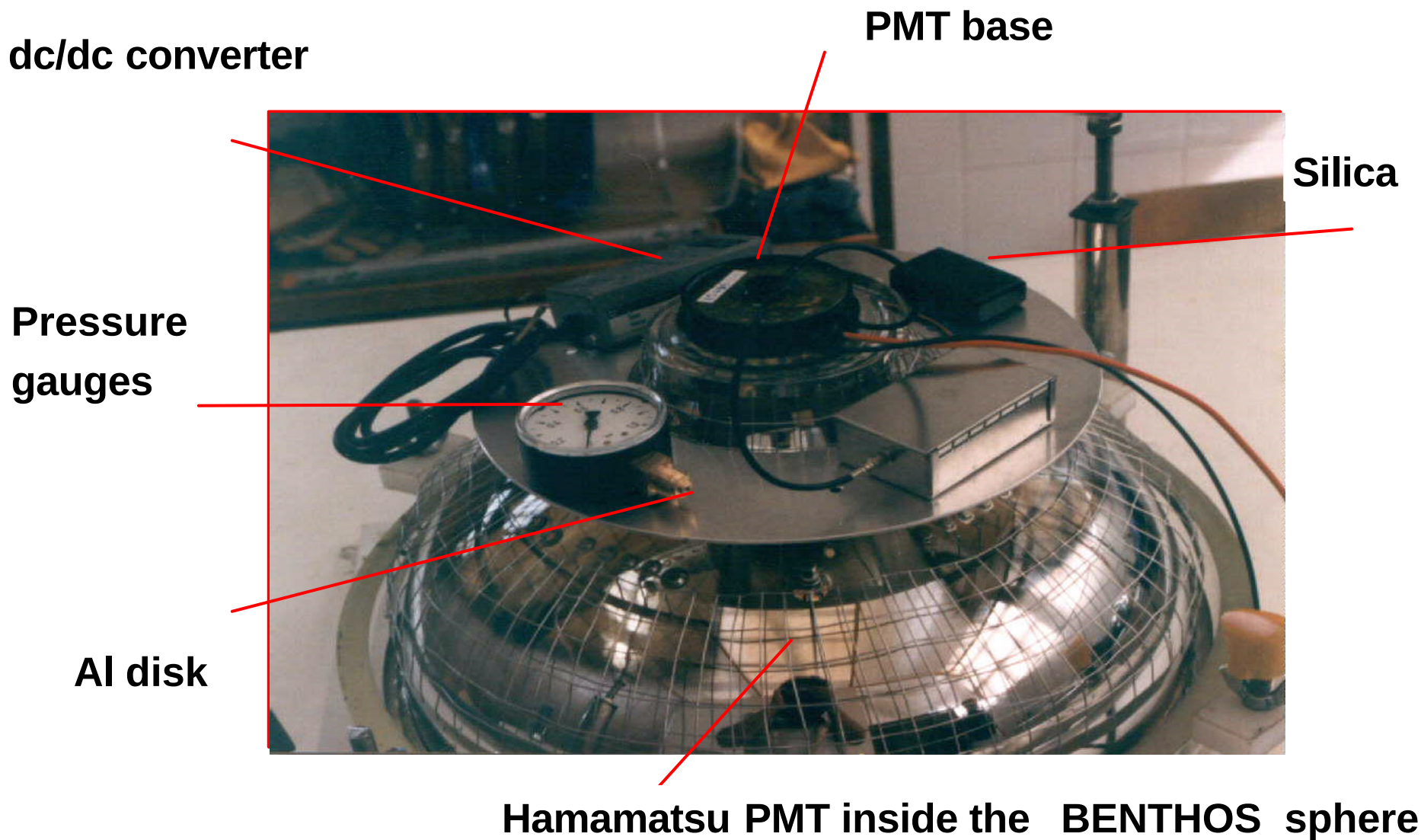


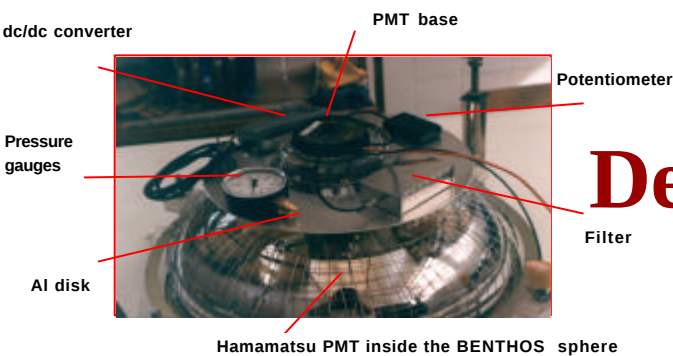
Filter

Hamamatsu PMT inside the BENTHOS sphere

- Hamamatsu PMT R2018-03 (15")
- Benthos spheres
- μ -metal cage
- power supply

The NESTOR Optical Module



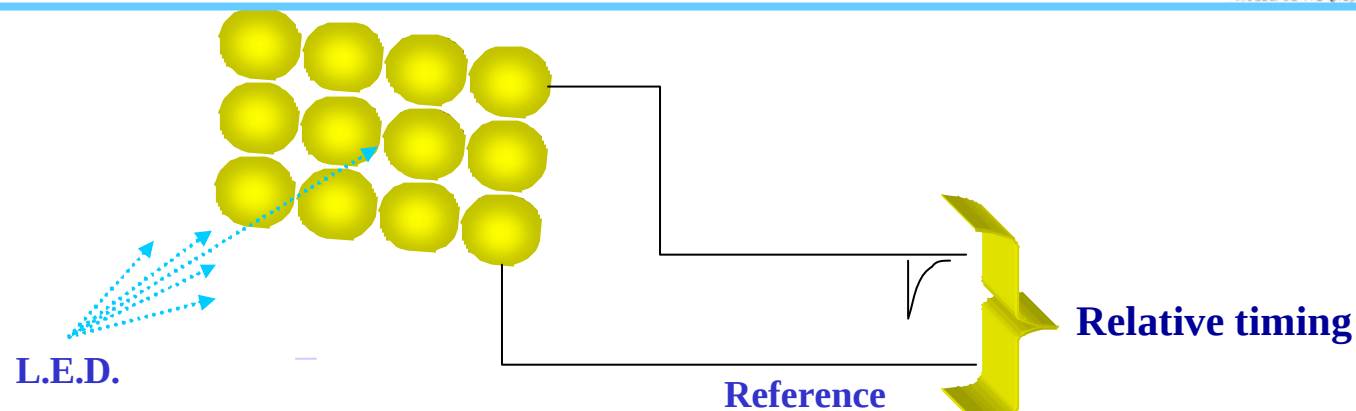
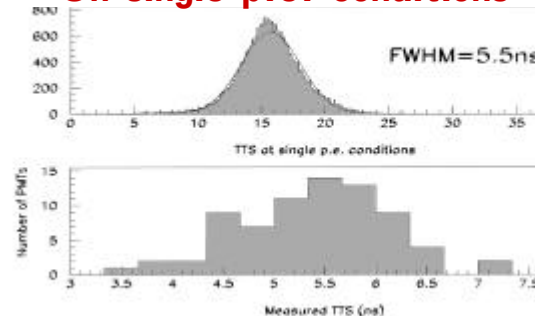


Detector Preparation

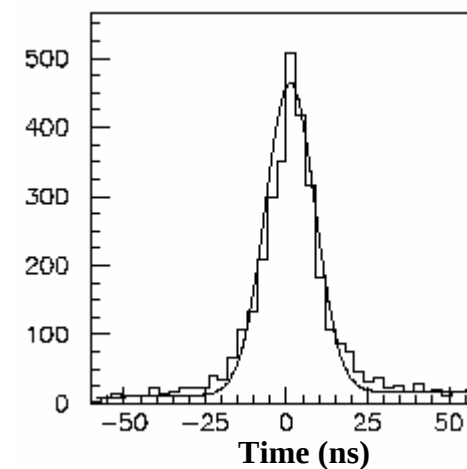
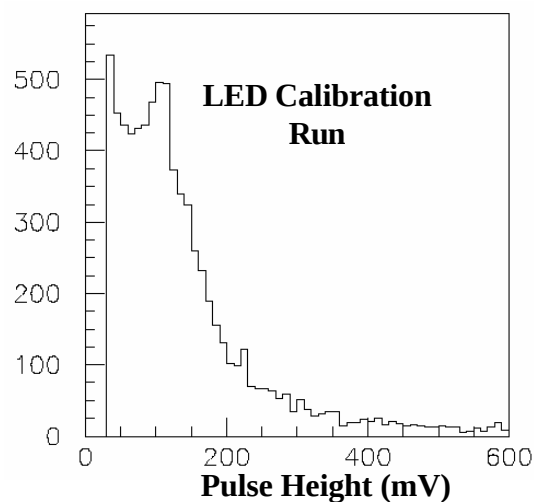
Optical Modules

Extensive Lab tests

On single p.e. conditions



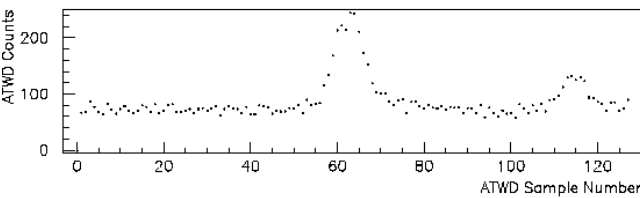
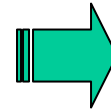
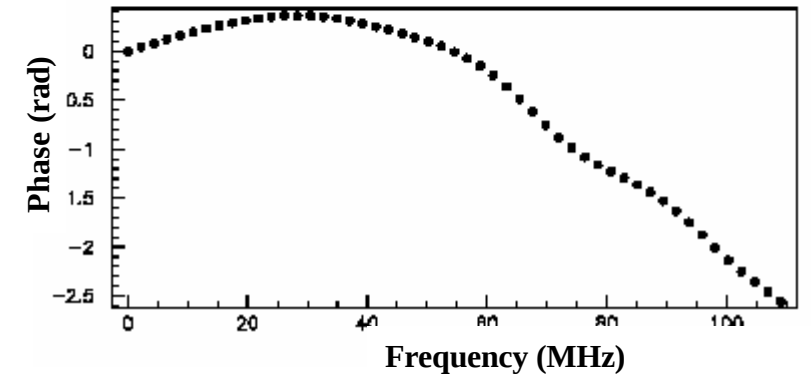
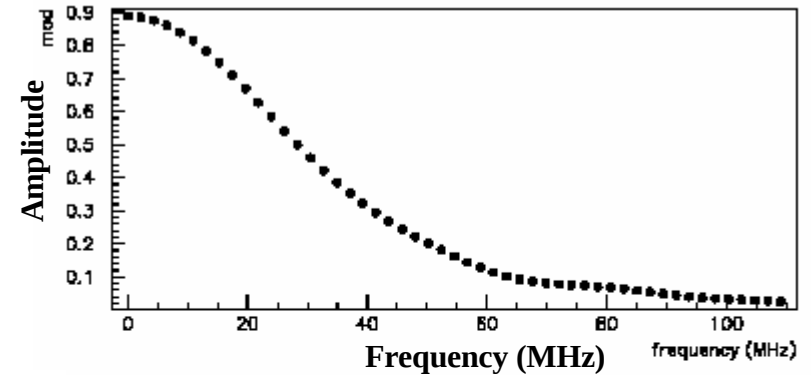
Time Difference in PMTs Response



Detector Preparation

Attenuation Correction

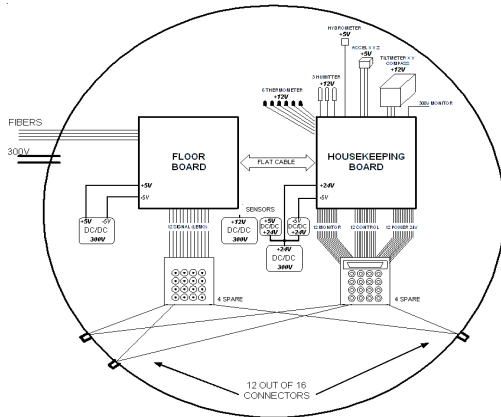
Fourier Transform-Comparison



Reference waveform

Electronic delay lines and amplifier

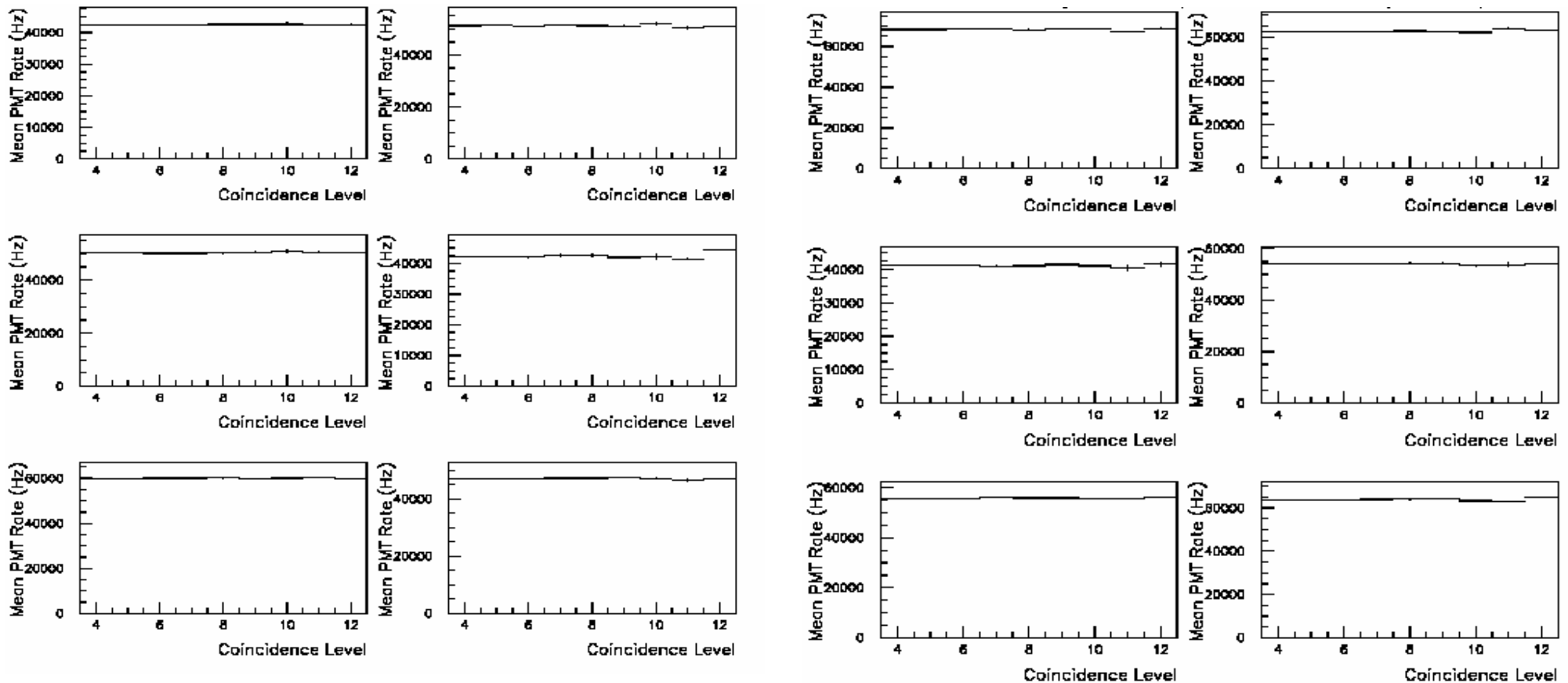
Coaxial cable



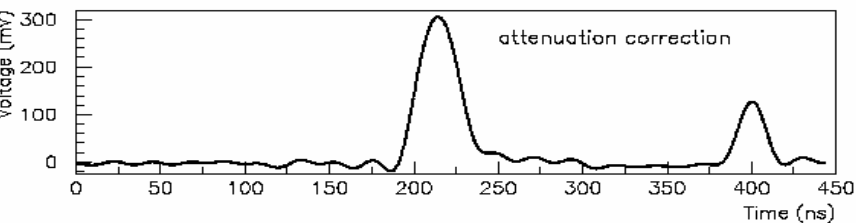
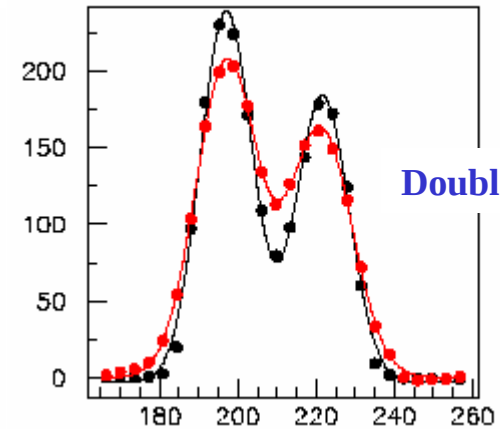
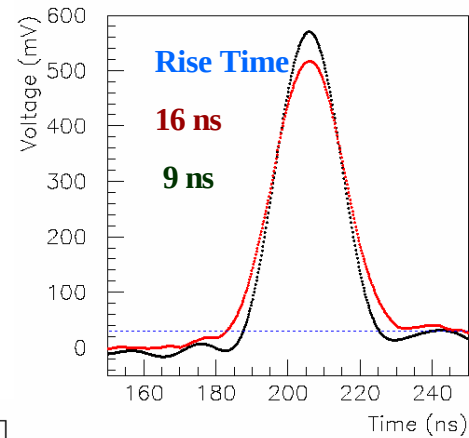
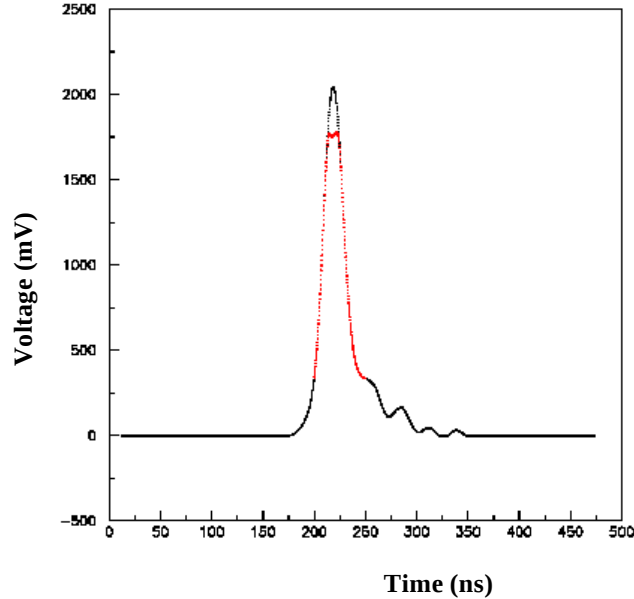
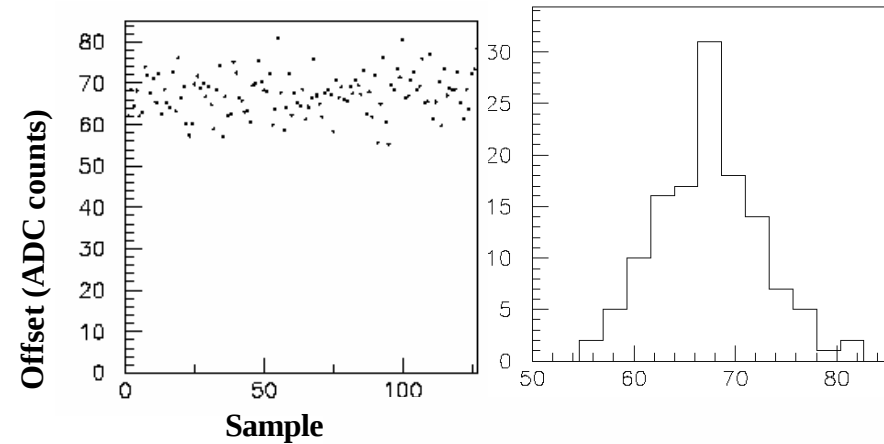
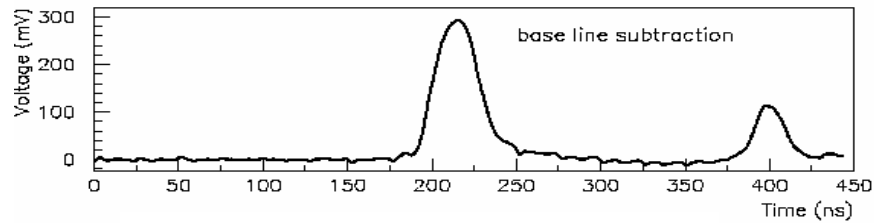
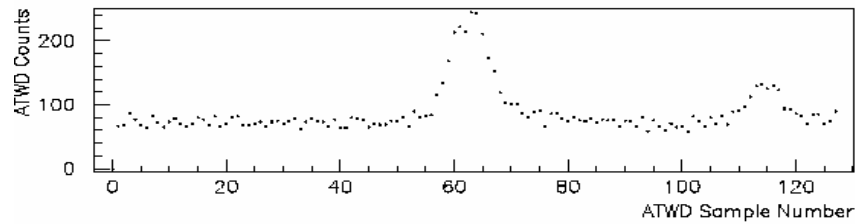
Data from a depth of 4000 m

Single PMT Rates

Trigger: =4fold Coincidence



Waveform Reconstruction

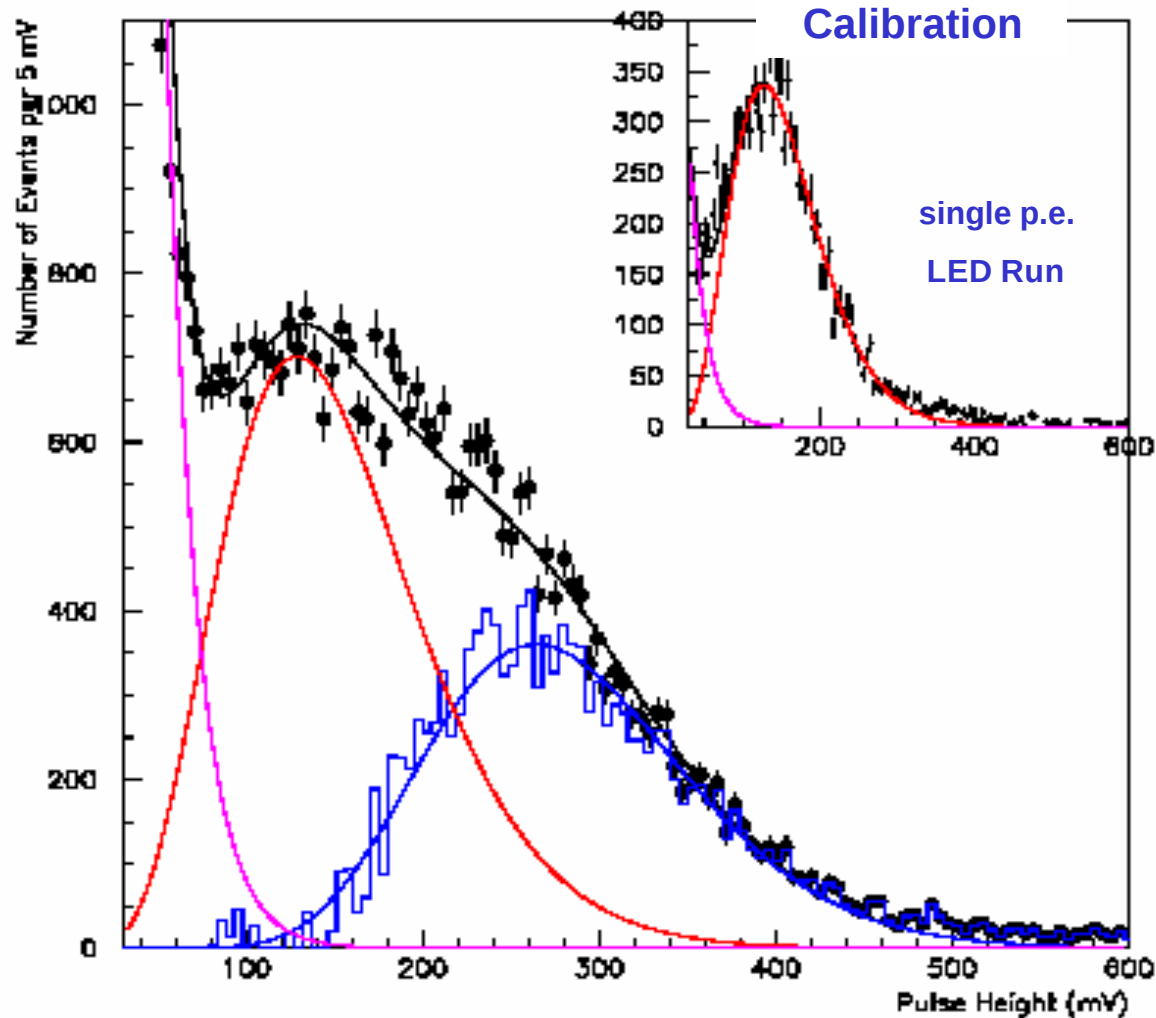


— Before the F.F.T. and the attenuation corrections

— After the F.F.T. and the attenuation corrections

Data from a depth of 4000 m

PMT Pulse Height Distribution

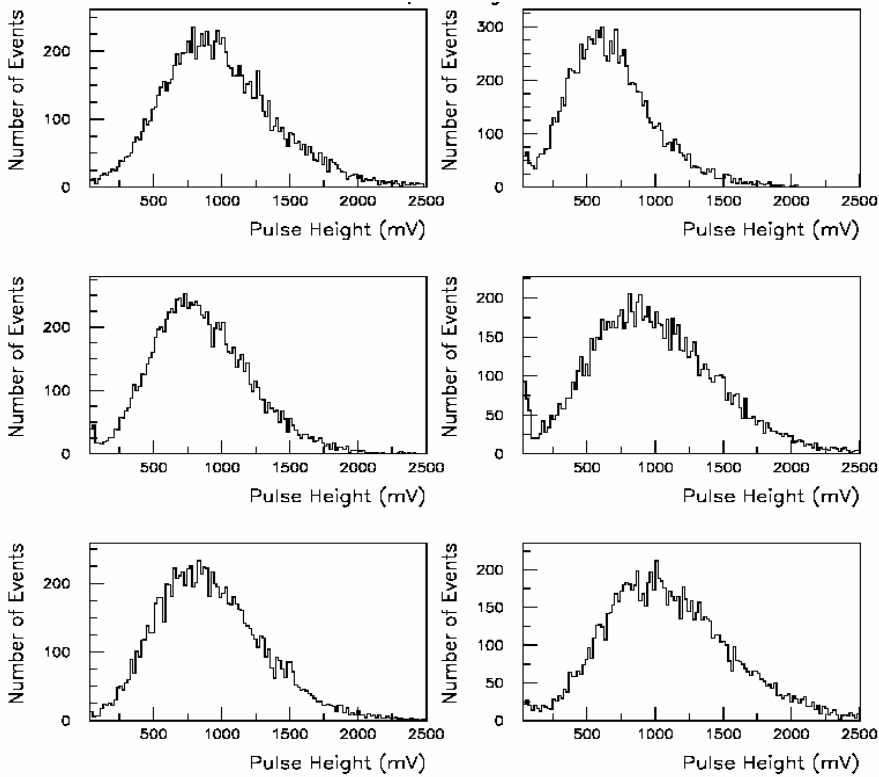


- single p.e. pulse height distribution
- two p.e.s pulse height distribution
- dark current pulse height distribution
- sum of the above

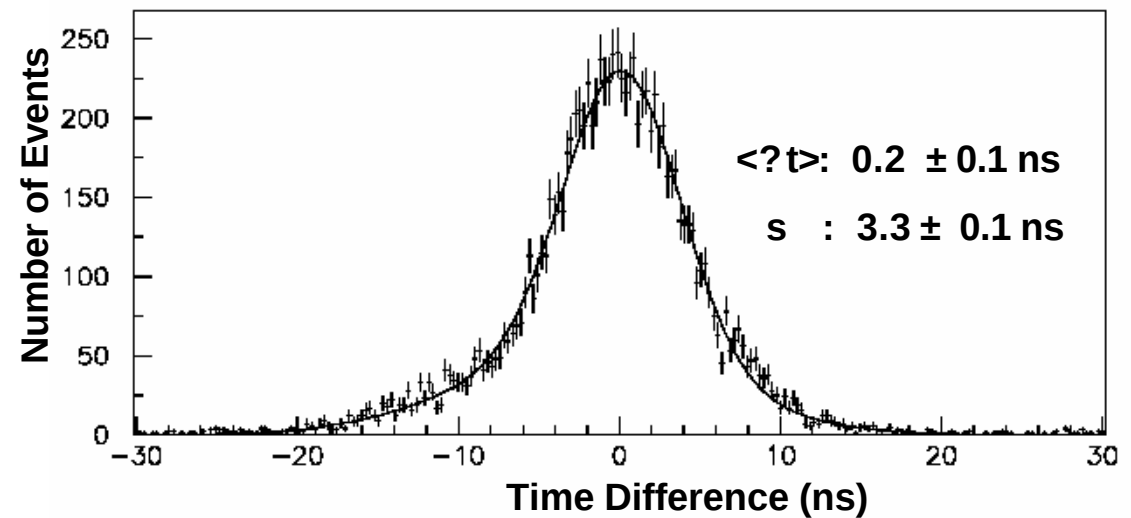
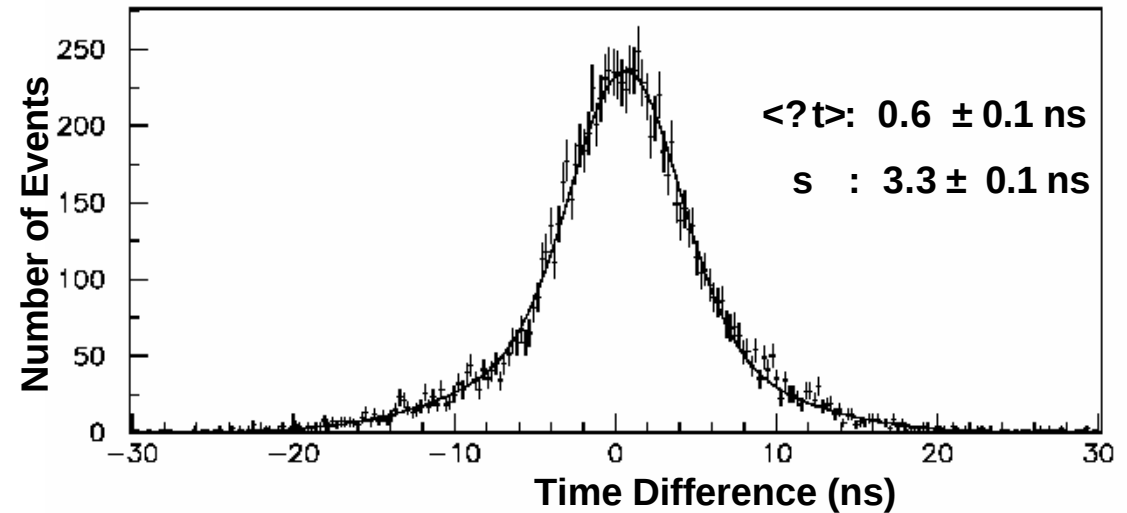


Calibration Data Analysis

Calibration Database
Quality Histograms



Data from a depth of 4000 m Calibration Run



Data from a depth of 4000 m

Calibration Run

